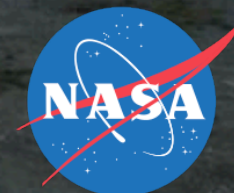


DYNAMO: Budget and TRMM Product Intercomparisons

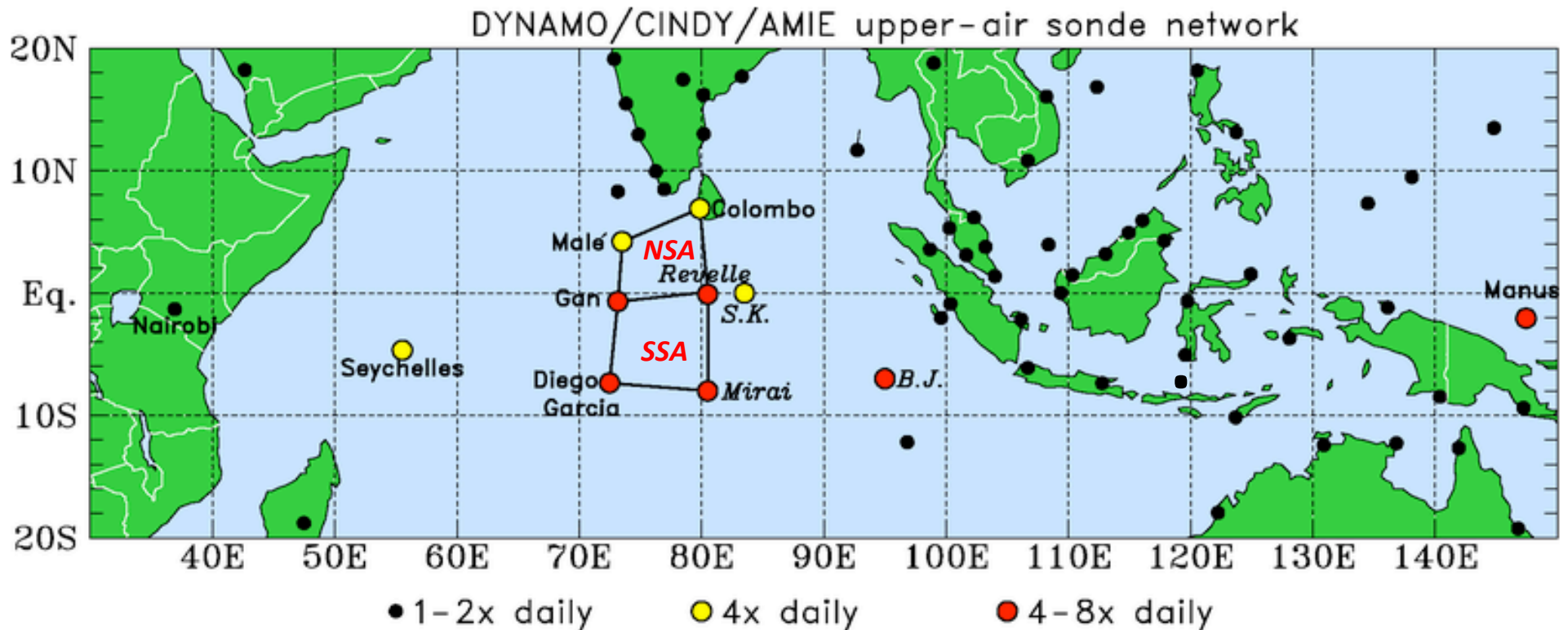
Richard H. Johnson and Paul E. Ciesielski
Colorado State University

Hanna Barnes photo

NASA Precipitation Measurement Missions (PMM) Science Team Meeting, 13-17 July 2015, Baltimore MD

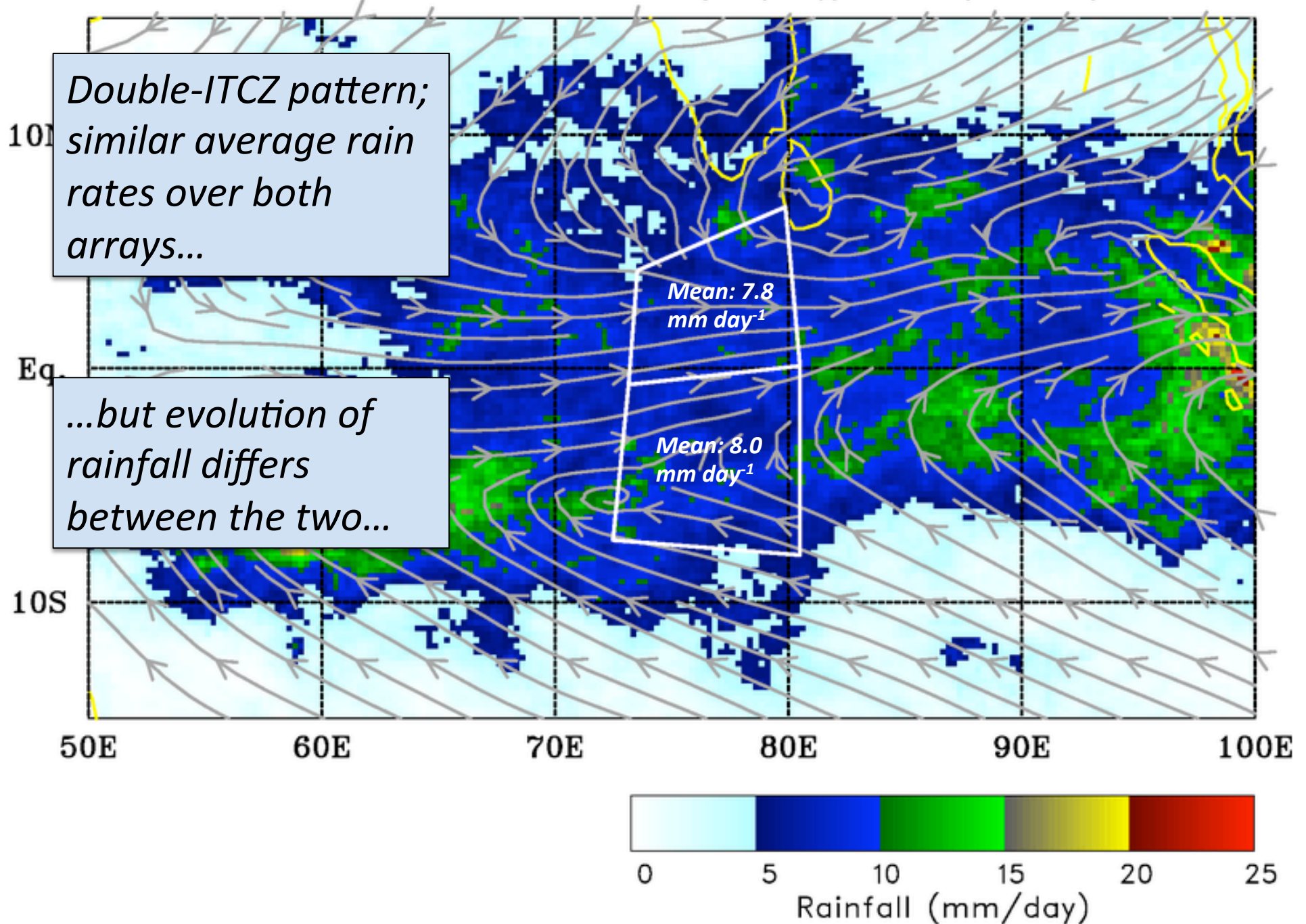


DYNAMO/CINDY/AMIE Sounding Network

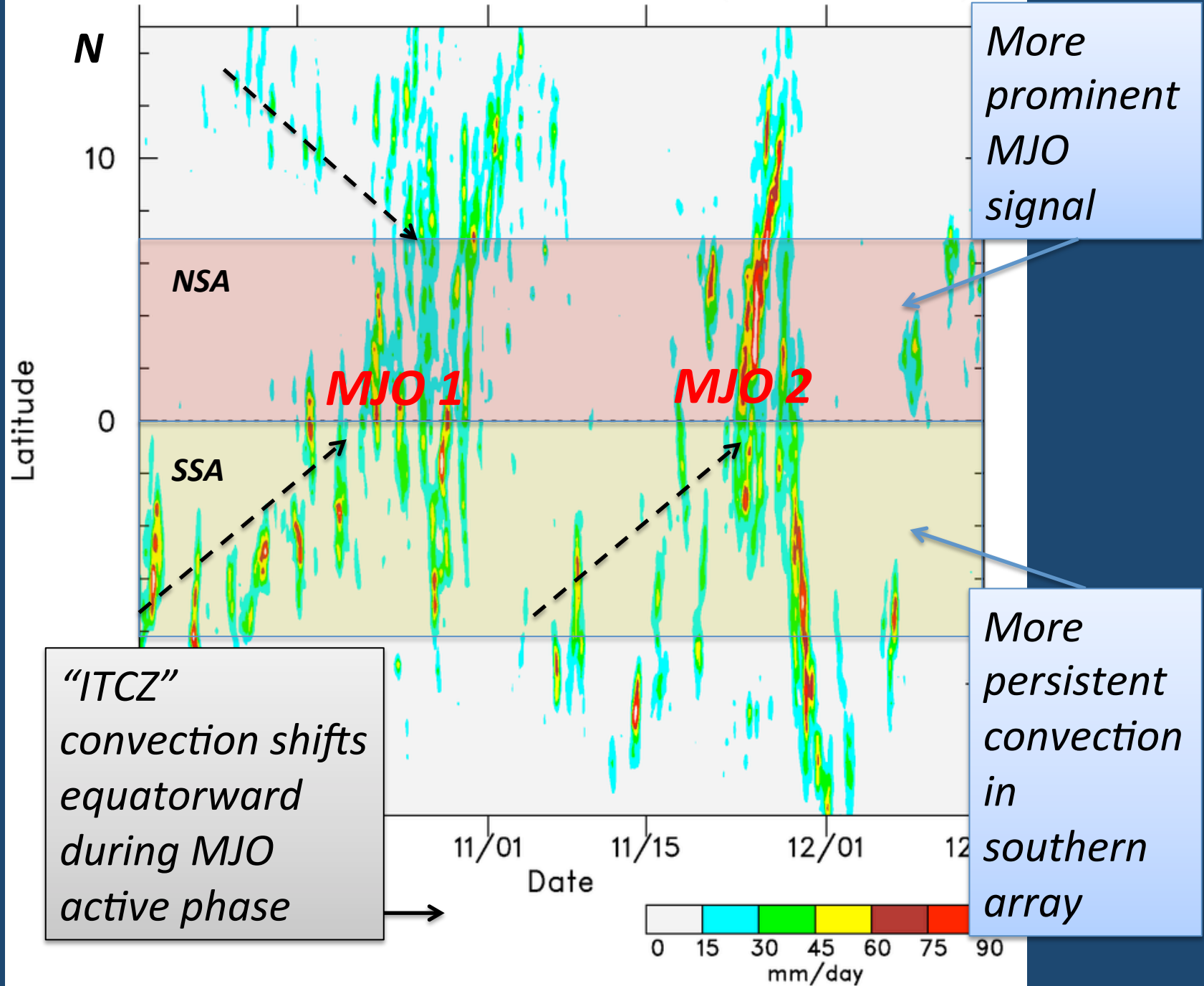


- **Objective:** investigation of the initiation of the MJO over the Indian Ocean – the roles of large-scale advection, cloud populations, and air-sea interaction
- Qc'ed sounding data, along with satellite winds and COSMIC data, were objectively analyzed onto a 1° grid; large-scale budgets were computed, averaged over sounding arrays and compared to TRMM products.

Surface streamlines and TRMM rainfall (mm/day) from 10/01 - 12/15 2011



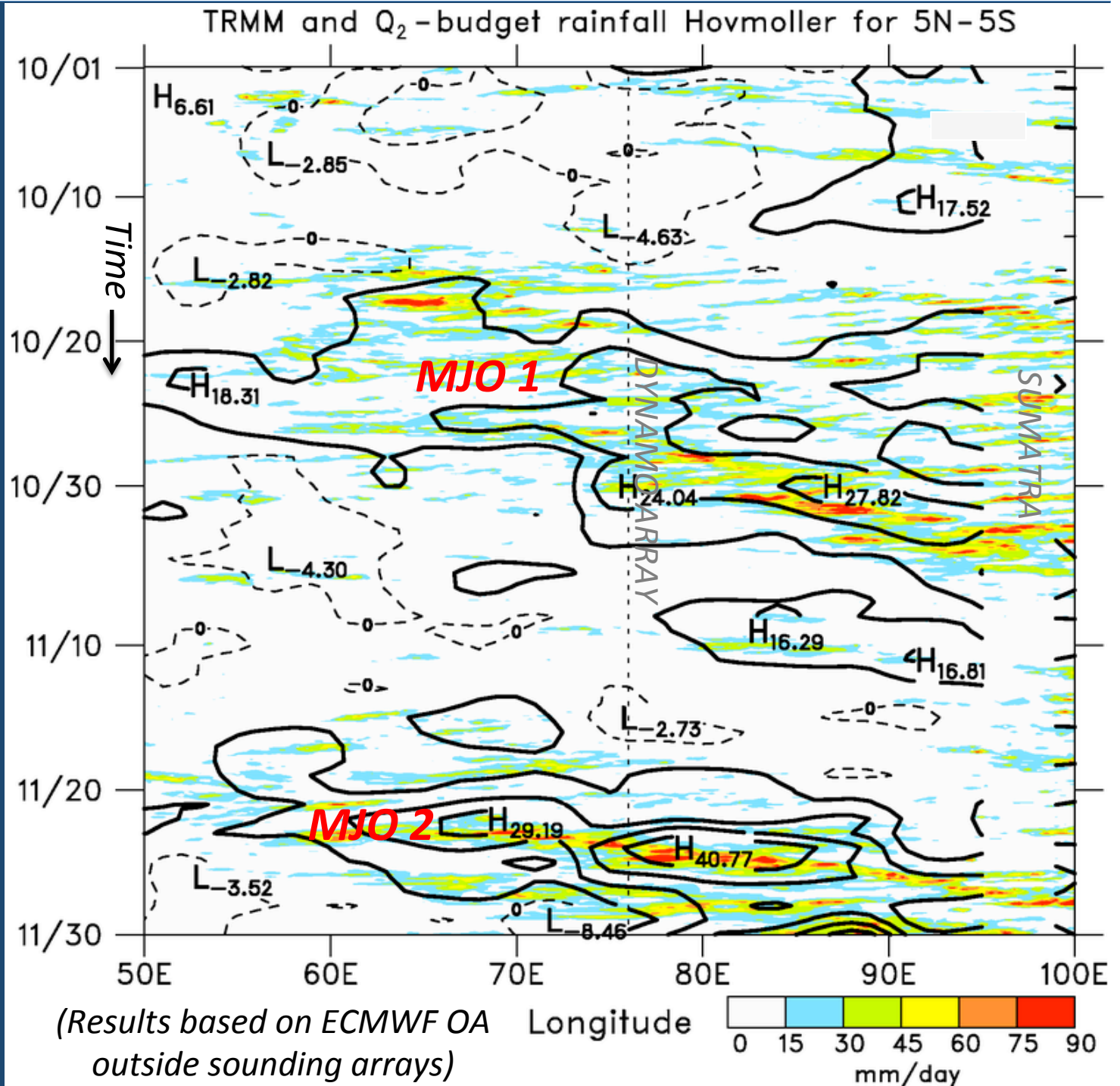
TRMM 3B42-V7 Hovmoller for 2011 (72.E to 80.E)



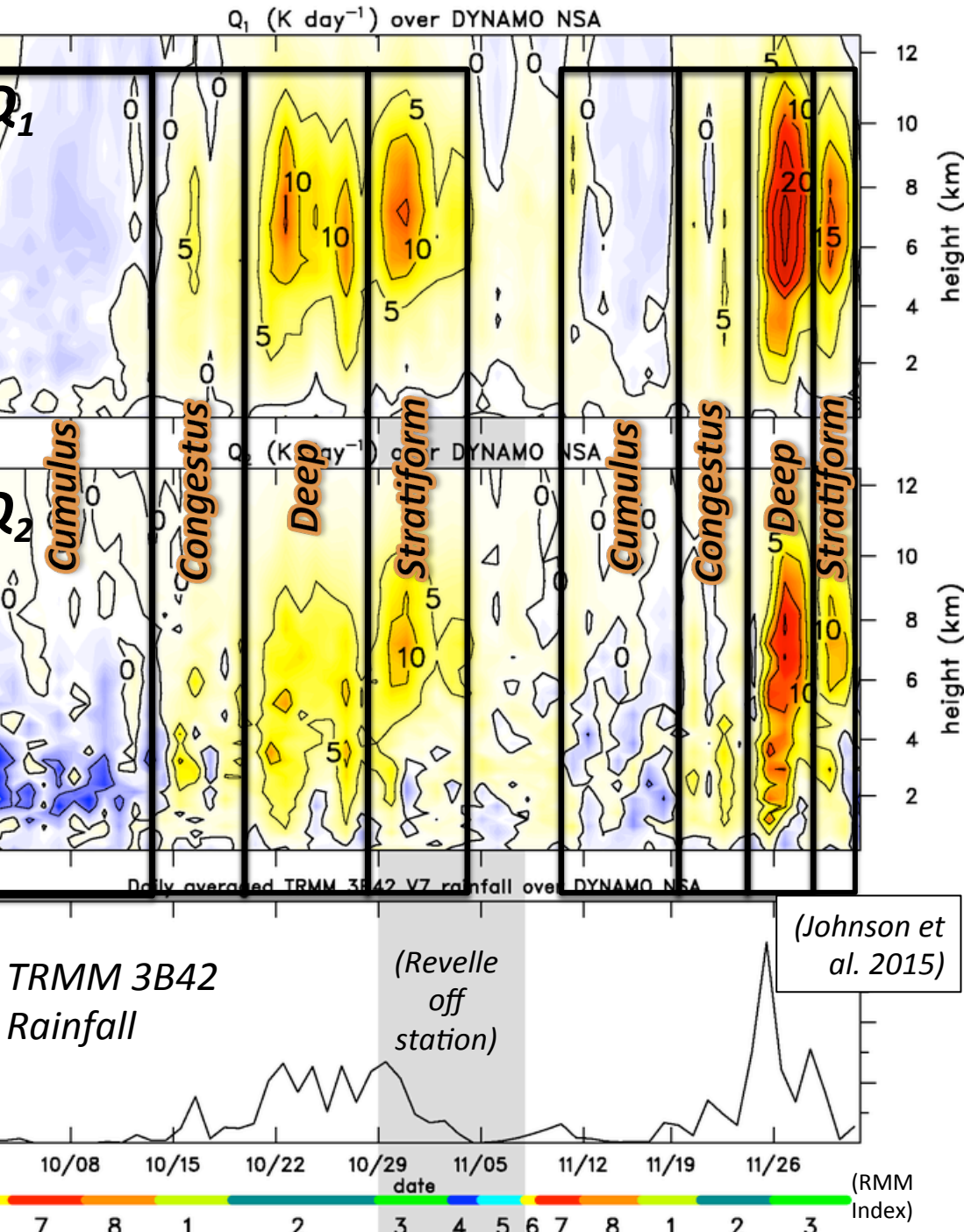
■ Two prominent MJO events: October and November

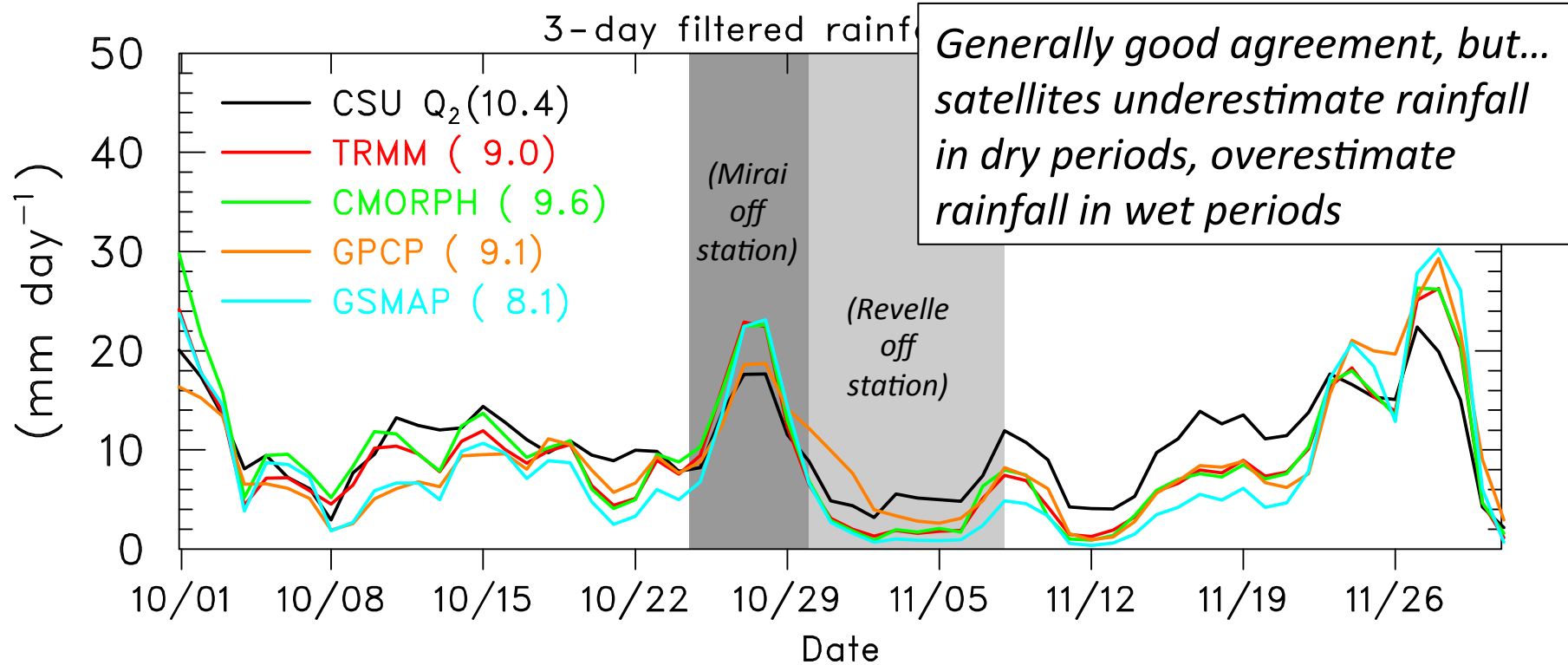
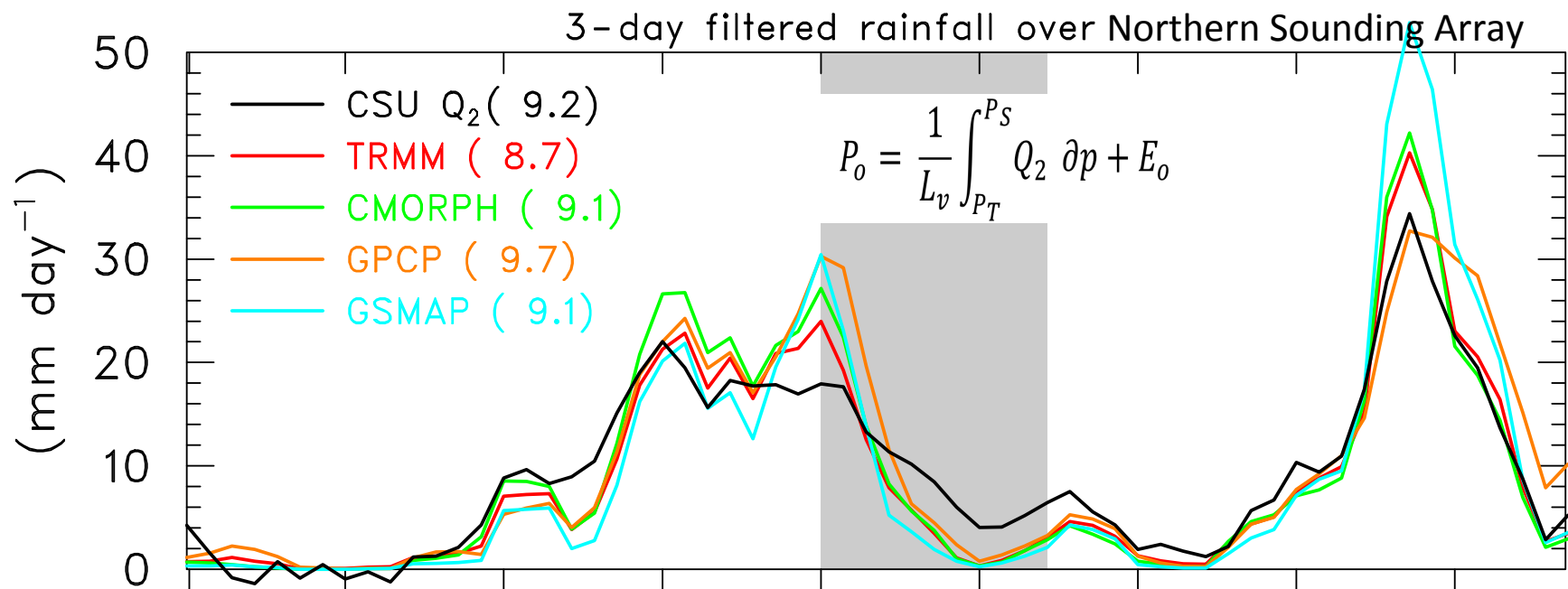
■ MJOs consist of westward- and eastward-moving disturbances

■ Q_2 budget captures precipitation envelopes

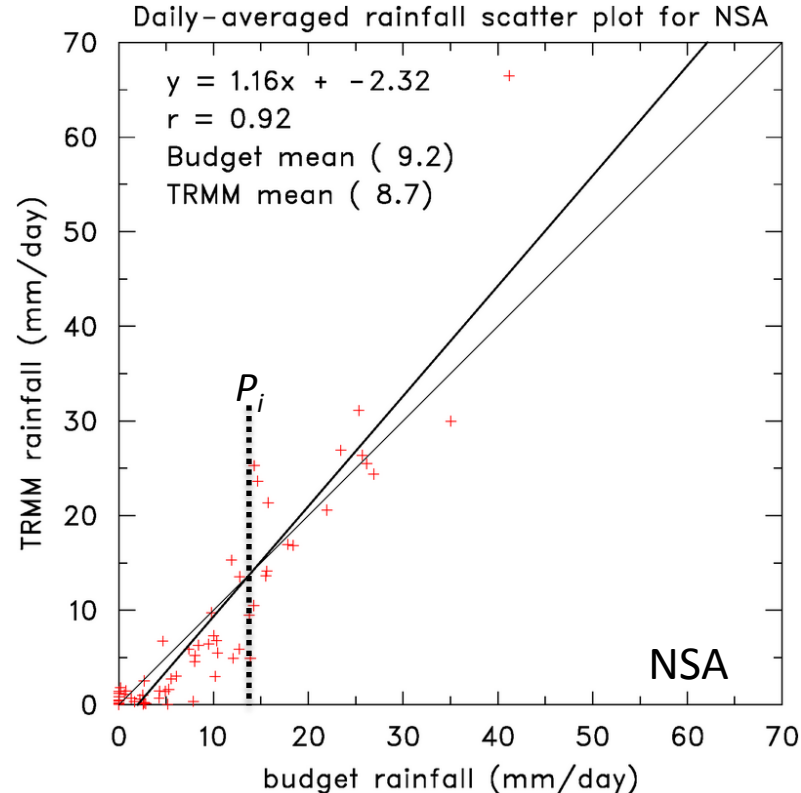


- Cumulus
- Congestus
- Deep convection
- Stratiform





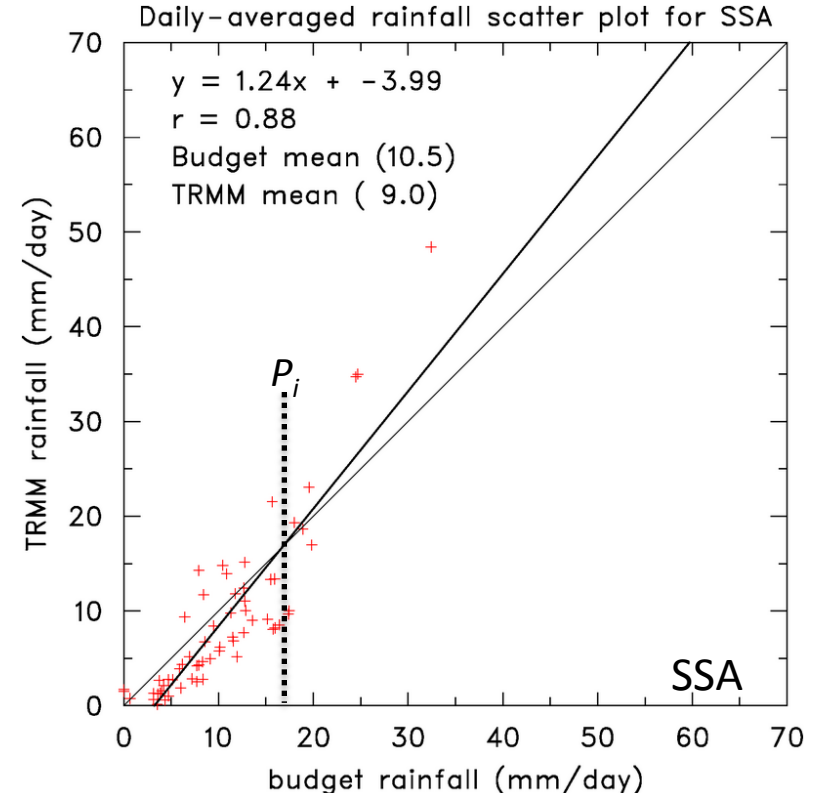
Scatter plots of Rainfall (TRMM 3B42 vs Budget-derived)



P_i @ 14.2 mm/day

Below P_i : TRMM – Budget = -2.1 mm

Above P_i : TRMM – Budget = 3.7 mm

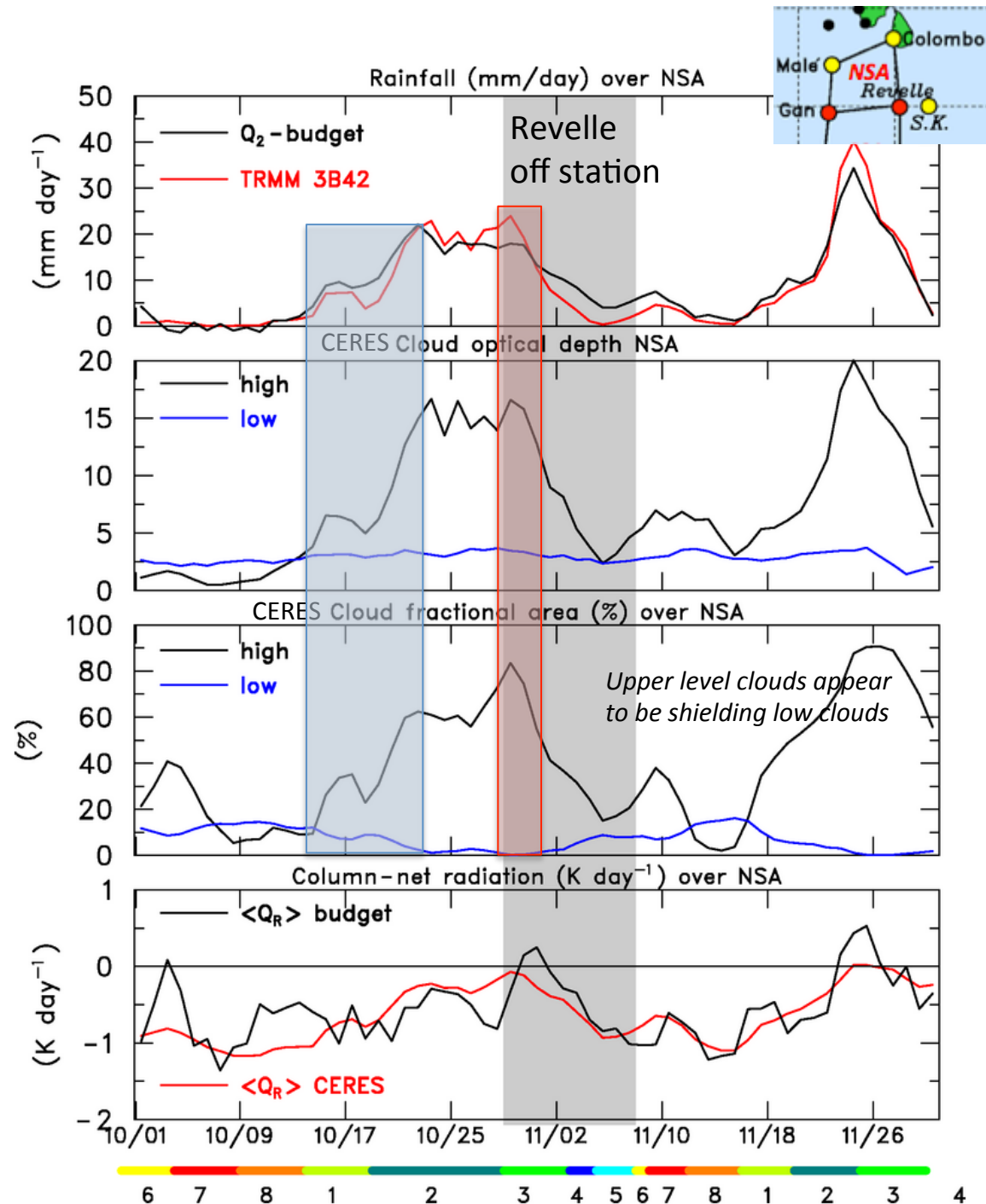


P_i @ 16.6 mm/day

Below P_i : TRMM – Budget = -2.5 mm

Above P_i : TRMM – Budget = 5.5 mm

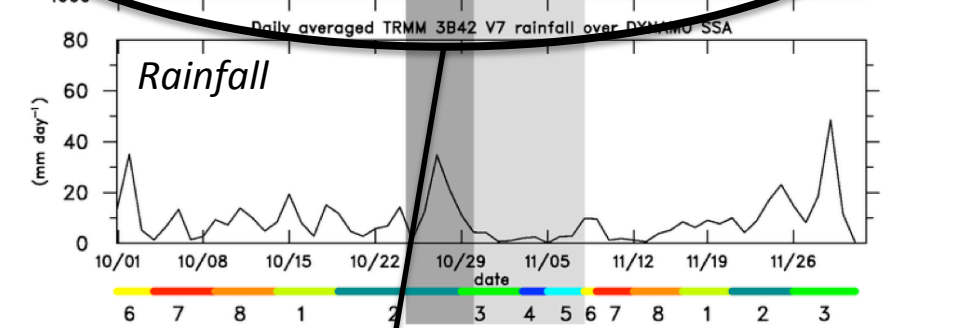
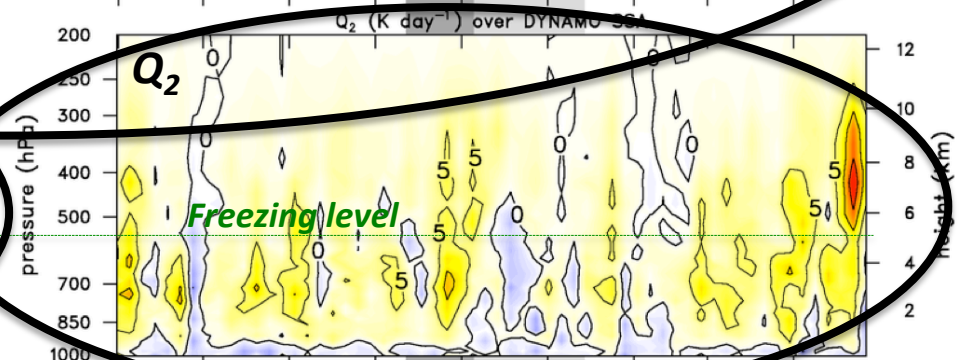
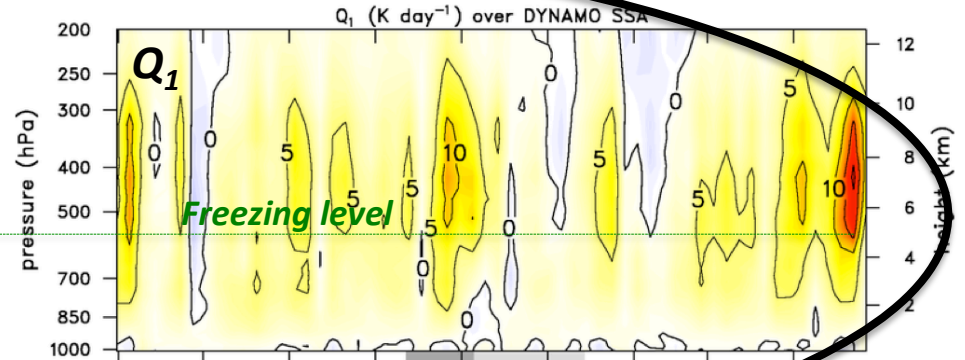
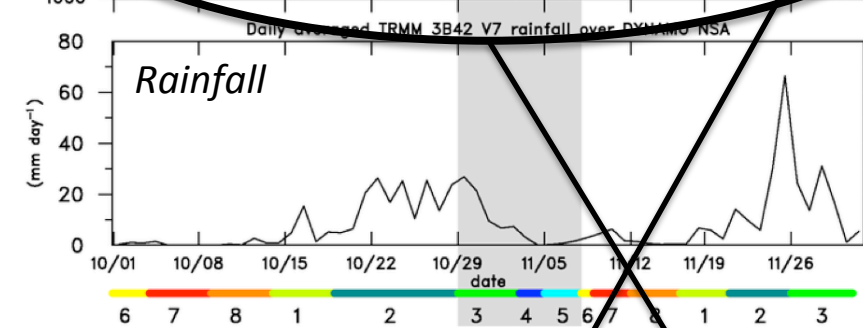
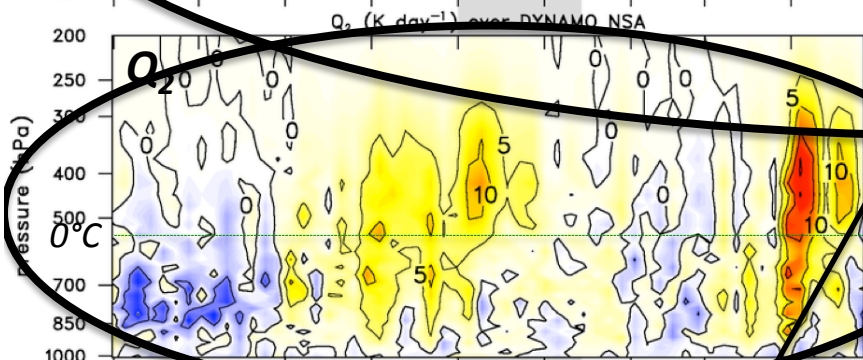
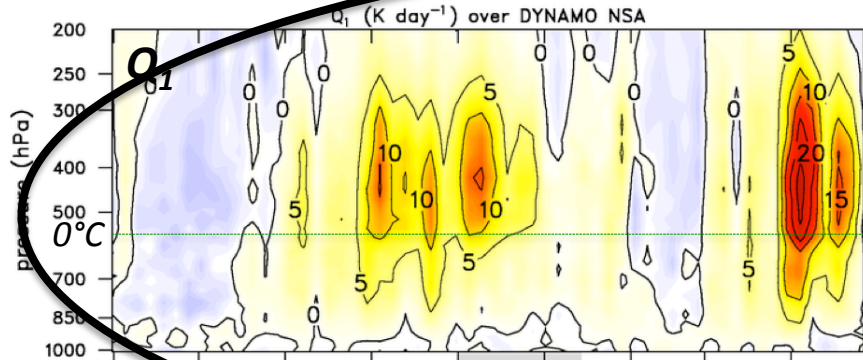
- This corroborates the finding of Xu and Rutledge (2014), based on radar data from RV Revelle, where they found that TRMM 3B42 rainfall product:
 - **underestimates** rainfall during suppressed periods (presumably due to insufficient sampling of shallow, warm-rain clouds)
 - **overestimates** rainfall during convectively active periods (likely due to abundance of high-level cloudiness).



- Rainfall differences increase over NSA when Revelle was off station
- Condensate storage in clouds may be contributing to rainfall differences.
- When cloud field is increasing, budgets will tend to overestimate rainfall
- Conversely, when cloud field is decreasing, budgets tend to underestimate rainfall
- Strong correlation between $\langle Q_R \rangle$ and upper-level cloudiness (i.e., $|\langle Q_R \rangle|$ decreases as high-cloudiness increases) confirms important role of high cloudiness in trapping longwave radiation.

NSA

SSA

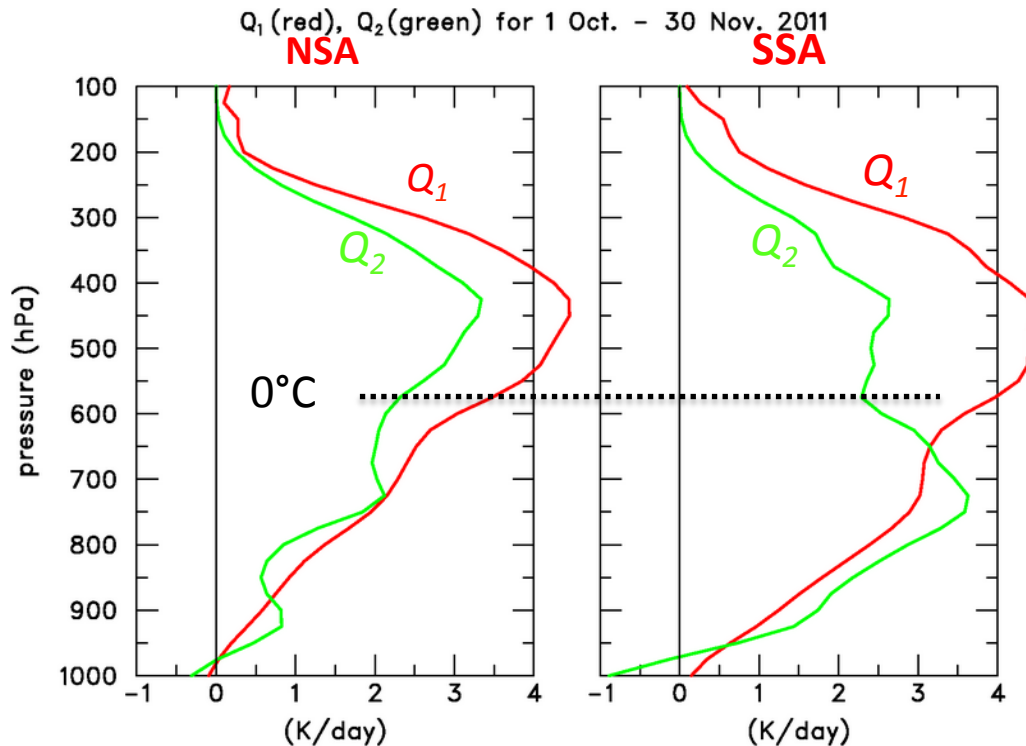


Most Q₁ peaks in upper troposphere

Q₂ peaks mostly in lower troposphere

Q₂ peaks in upper and lower troposphere

2 month-mean vertical profiles



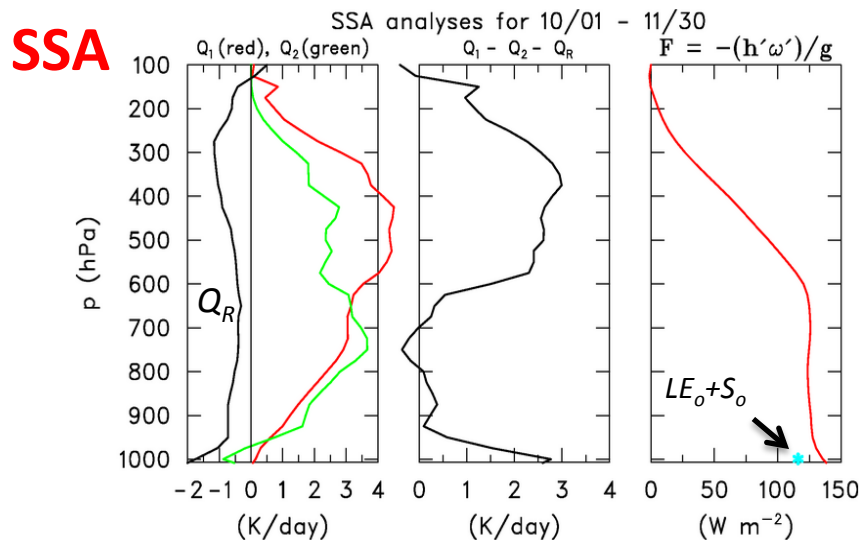
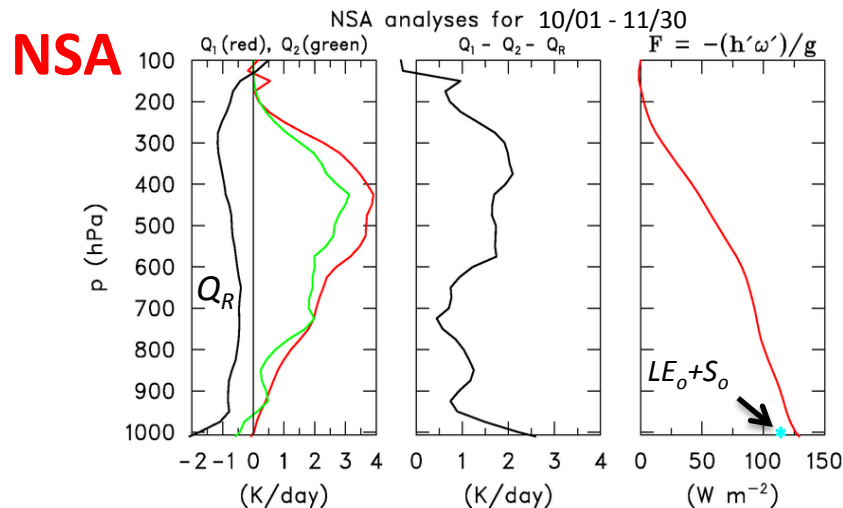
- Both arrays have peak heating level near 425 hPa; peak for SSA is somewhat broader.
- Q₂ peaks near 400 hPa for NSA; below 700 hPa for SSA

- Over the NSA, Q₁ and Q₂ peaks are coincident in the vertical; while over the SSA a large vertical separation is present in these peaks
- Suggest that vertical convective eddies are stronger over SSA

Profiles of vertical eddy flux of moist static energy

$$F(p) = -\frac{1}{g} (h' \omega') = \frac{1}{g} \int_{p_T}^p (Q_1 - Q_2 - Q_R) \partial p$$

- $Q_R(p)$ over arrays is estimated by adjusting Gan $Q_R(p)$ so that its vertical integral matches the CERES array-averaged $\langle Q_R \rangle$



- Upper-level eddy flux convergence is larger over SSA resulting in stronger and deeper convective eddy fluxes (F) over this region.
- Weaker F over NSA implies that stratiform rainfall fraction (SF) may be more prominent over NSA.
- During this period SF from TRMM PR 2A25 is 55% over NSA and 50% over SSA for this period (*M. Katsumata*).

Validation of TRMM LH products for DYNAMO Oct-Nov. 2011 period

CSH (Convective-Stratiform Heating) :

- *Tao et al. (1993, 2000, 2001, 2006, 2010, 2015)*
- *Developers: Tao and Lang*
- 3H31v7
- **Monthly**, 0.5°, 19 vertical levels (0.5, 1, 2, ... 18km)
- Q_1, LH, Q_R, Q_2

SLH (Spectral Latent Heating)

- *Shige et al. (2004, 2007, 2008, 2009)*
- *Developers: Shige and Takayabu*
- 3H25v7
- **Monthly**, 0.5°, 19 vertical levels
- $Q_1 - Q_R, LH, Q_2$

$$Q_1 - Q_R = LH + \text{vertical eddy heat divergence}$$

$$Q_2 = Q_{c-e} + \text{vertical eddy moisture divergence}$$

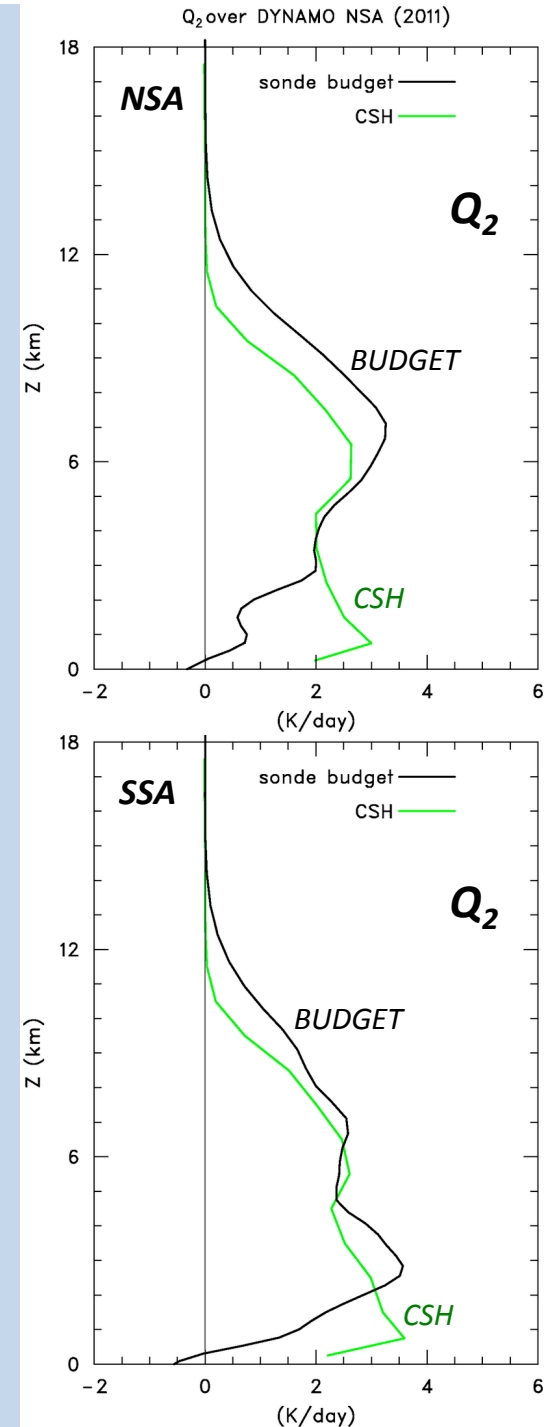
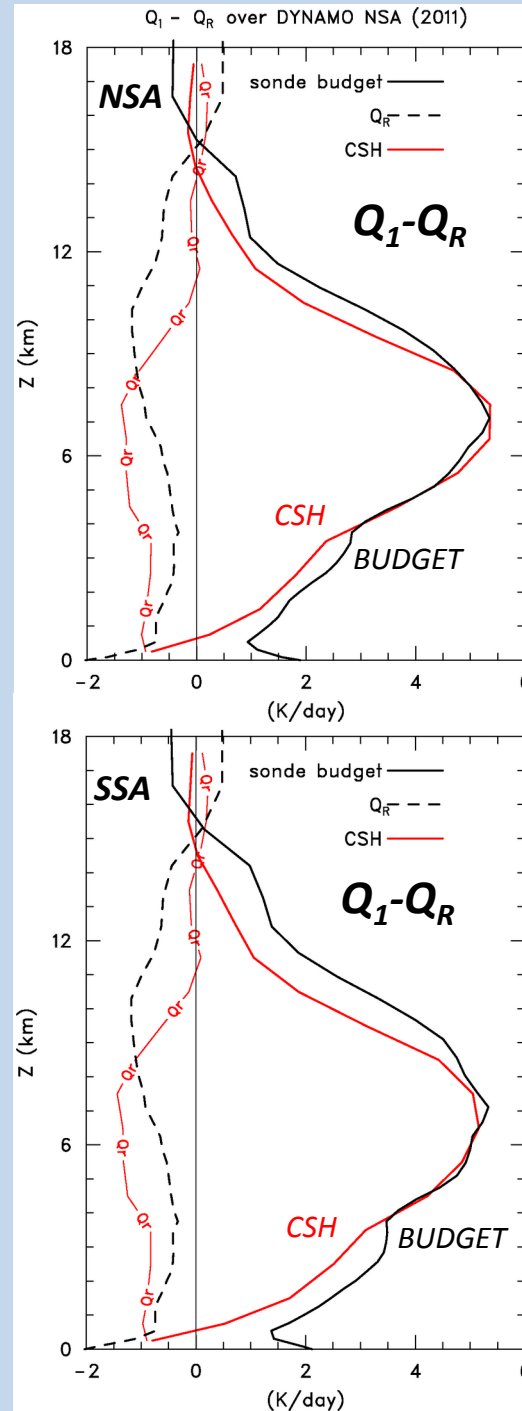
DYNAMO NSA/SSA Profiles:

Budget vs. CSH

Oct.-Nov. 2011

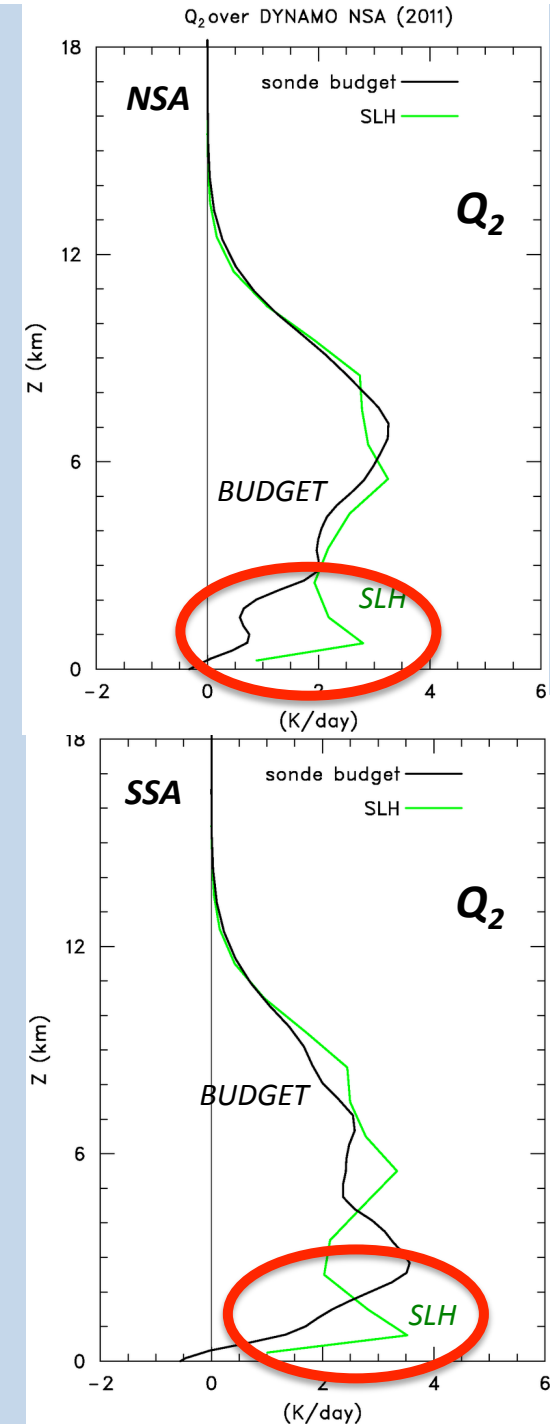
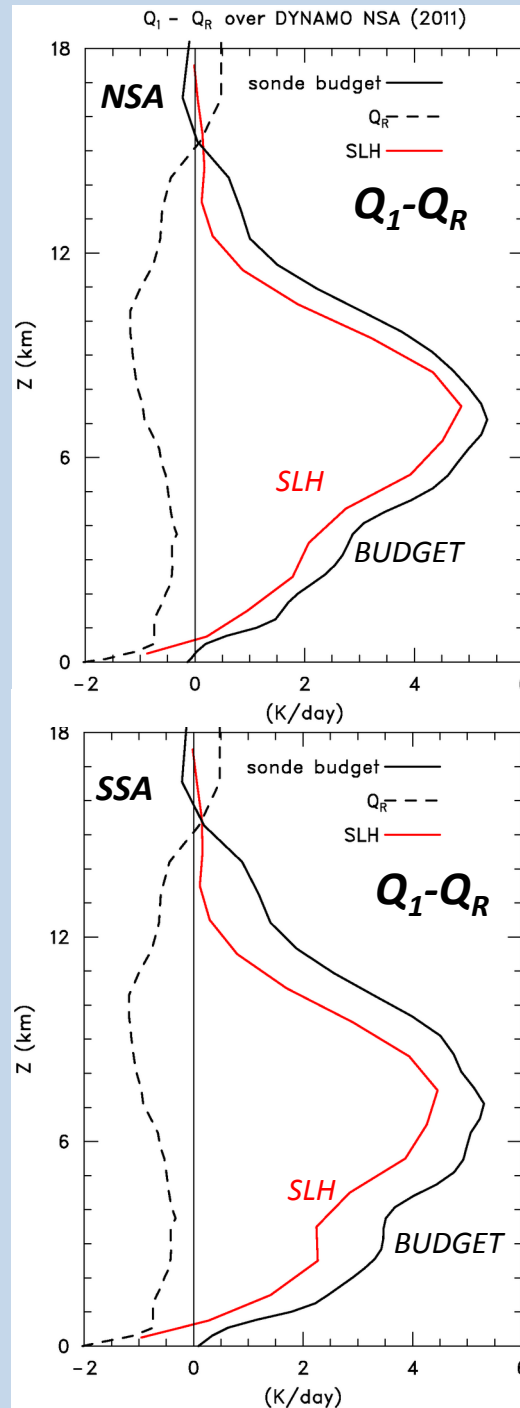
- $Q_1 - Q_R$ budget profiles show good agreement with CSH
- Q_R differences account for some of the discrepancy in $Q_1 - Q_R$
- Budget Q_2 profiles differ between N and S arrays; CSH has similar Q_2 profile for both arrays
- Too much low-level drying in CSH

(Q_R from Feng et al. 2014 based on Gan and CERES observations)

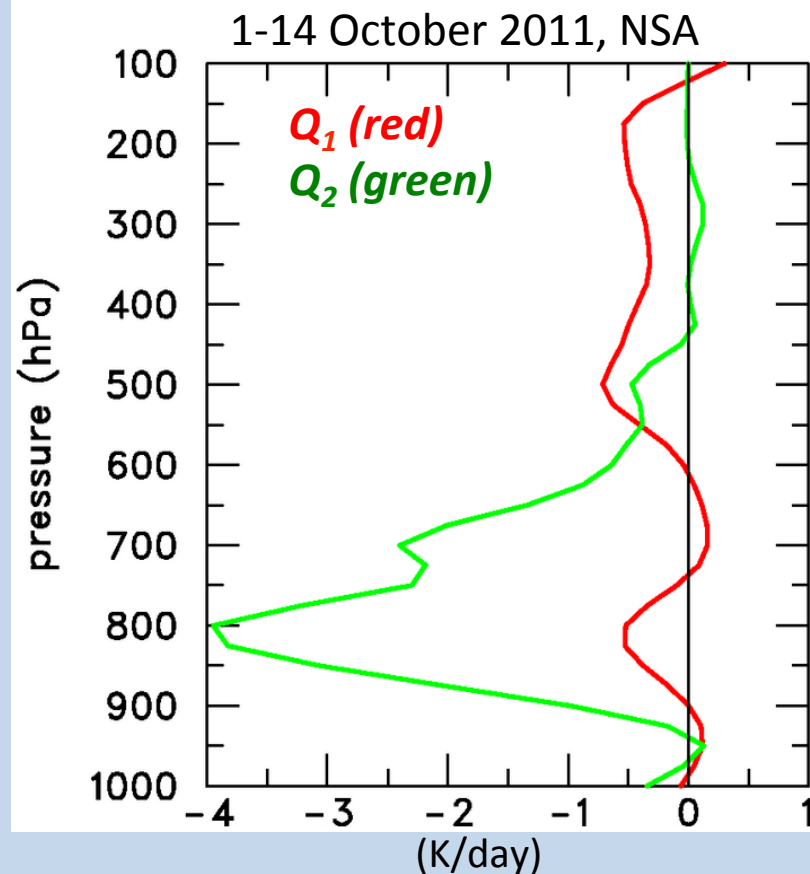
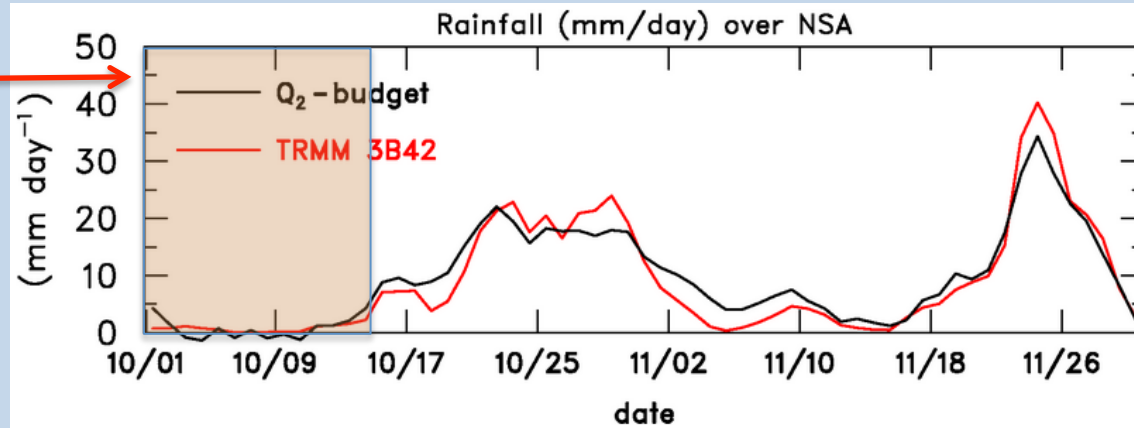


DYNAMO NSA/SSA Profiles: **Budget vs. SLH** **Oct.-Nov. 2011**

- SLH $Q_1 - Q_R$ amplitude about 10% less than budgets
- SLH (as for CSH) has similar Q_2 profile for both arrays; implies similar cloud populations whereas budgets do not
- Similar to the CSH, SLH has excessive low-level drying



- Little to no rainfall occurred during convectively suppressed period prior to October MJO active phase



- Strong low-level moistening and slight cooling during 1-14 October suggest the presence of shallow, non-precipitating convection during this period.
- Sensitivity of TRMM instruments prevent this type of convection from being detected which likely explains the excessive low-level drying in satellite Q_2 estimates.

Summary

- Overall good agreement of budget-derived and satellite-based rainfall estimates during DYNAMO
- Differences in rainfall are due to sampling issues (soundings and satellites) and hydrometeor storage effects
- Convection over NSA was strongly modulated by MJO signal, while the SSA experienced more persistent, briefer episodes of rain related to ITCZ convection.
- Weaker vertical eddy fluxes over the NSA compared to SSA and TRMM 2A25 analyses suggest a higher stratiform rainfall fraction over the NSA (Lin et al. 2004).
- CSH and SLH heating profiles are in good agreement with budgets but satellite-based Q_2 profiles show more significant differences particularly at low-levels.

A photograph of Dick Johnson, a man in a light-colored short-sleeved shirt and trousers, standing at night and holding a large, illuminated white balloon. The balloon has the words "Thank you" written on it in a stylized font. The background is dark with some distant lights and structures visible.

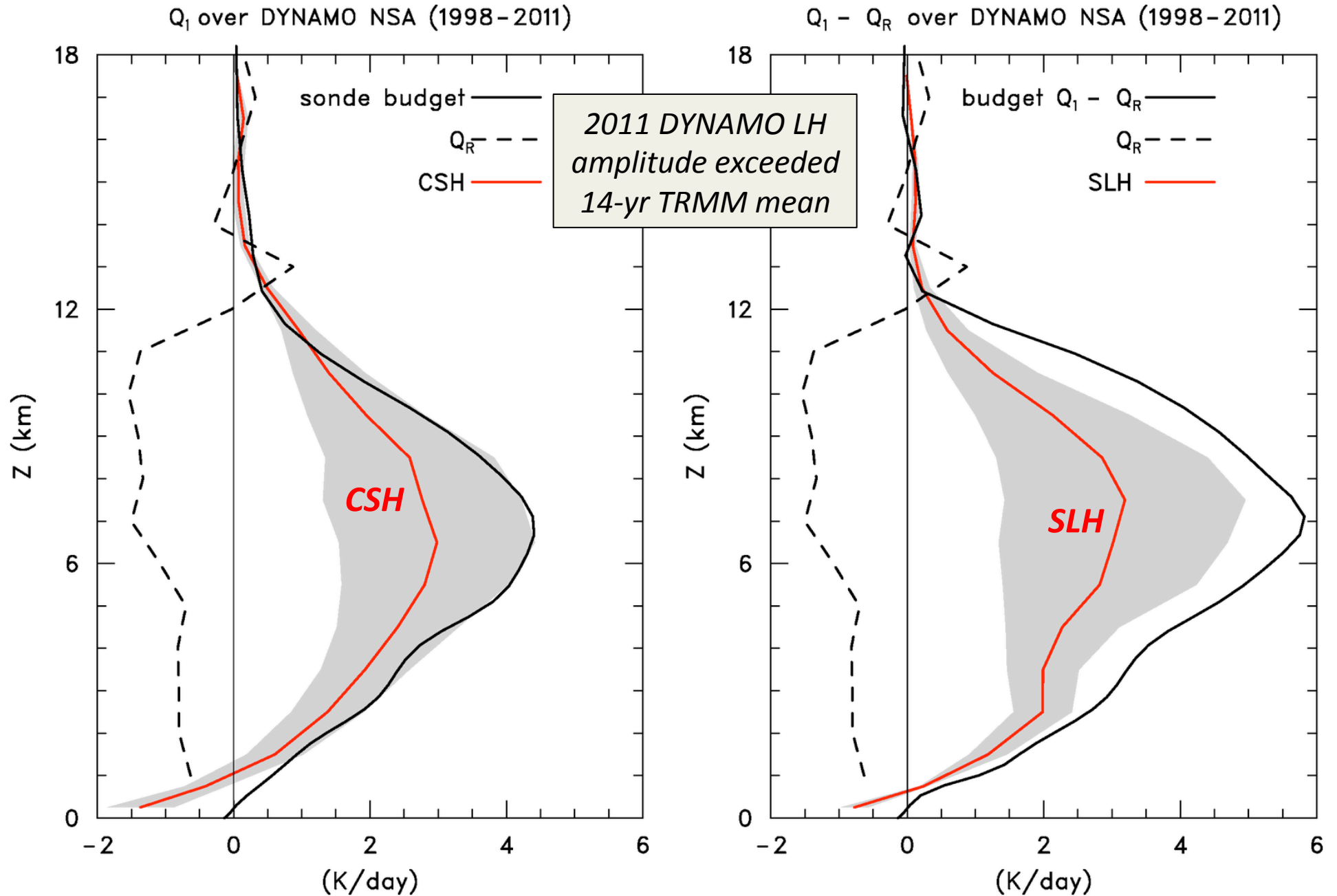
Thank you

Dick Johnson launching a nighttime sounding in Malé during DYNAMO

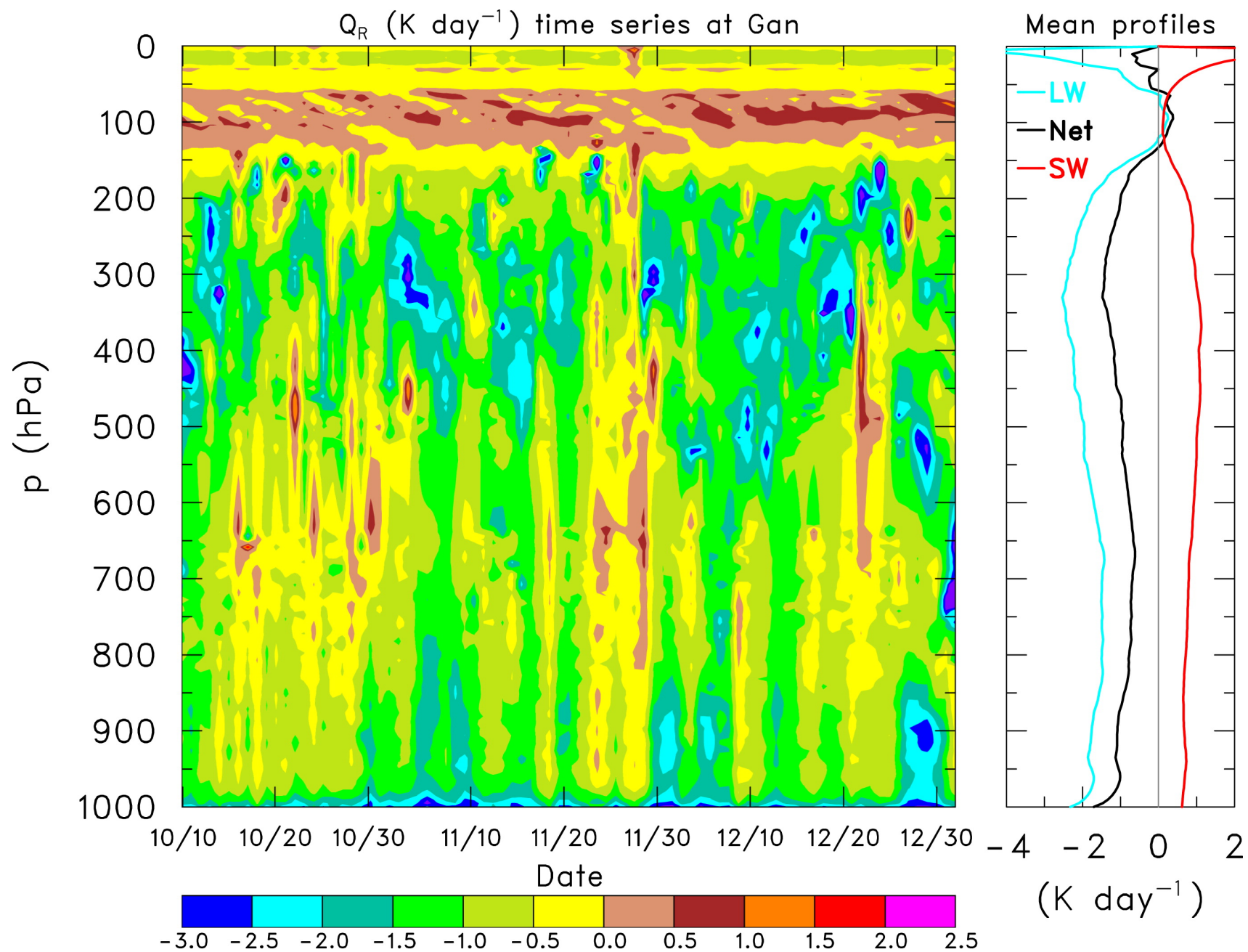
Supplemental material to follow

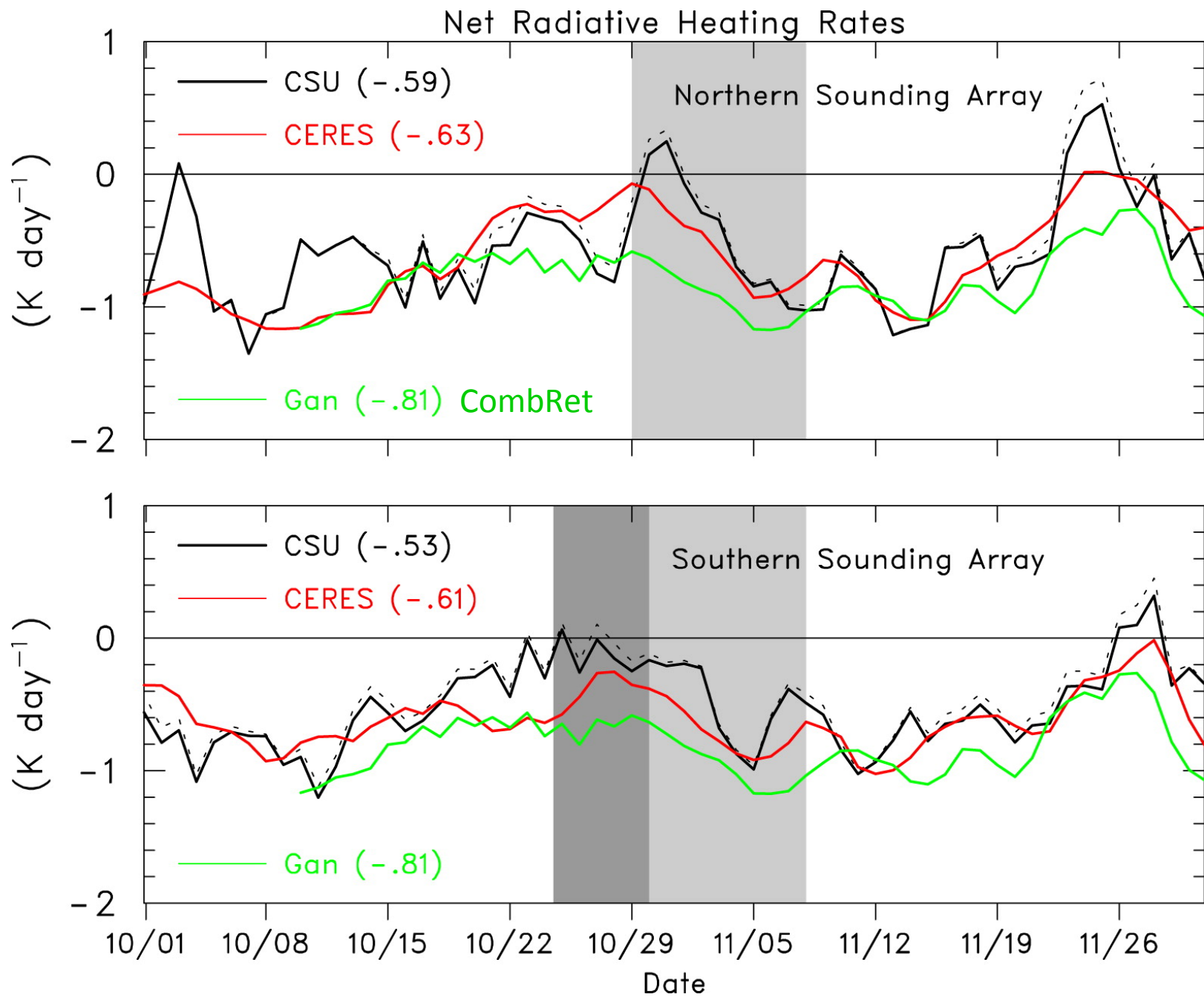
CSH (3H25): *Tao et al. (1993, 2000, 2001, 2006, 2010)*

SLH (3H31): *Shige et al. (2004, 2007, 2008, 2009)*

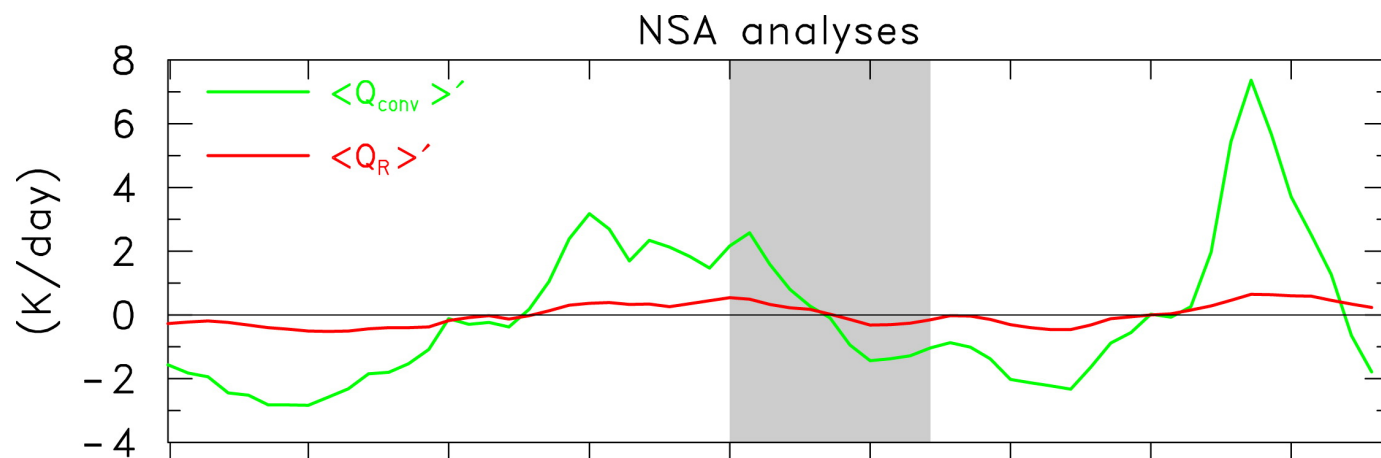


Radiative Heating Rates at Gan Island from CombRet

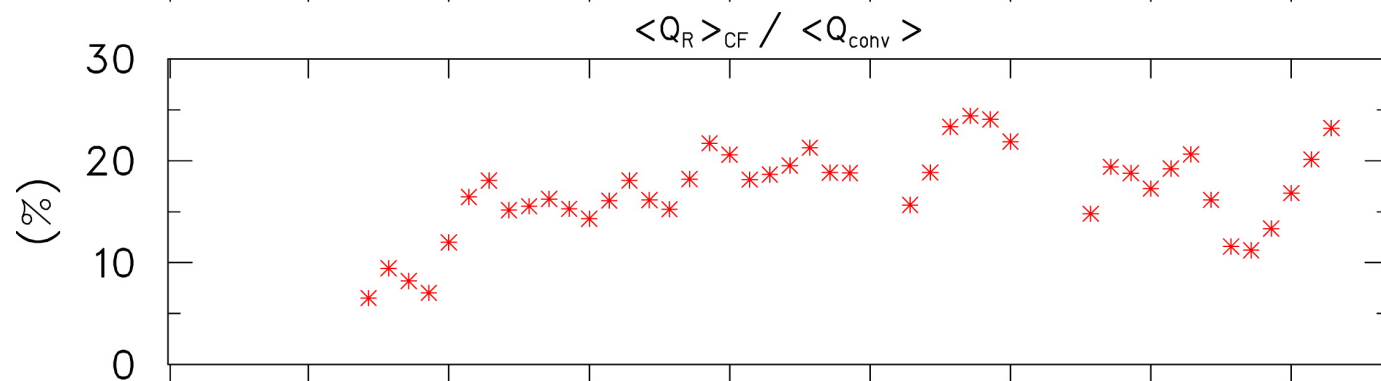




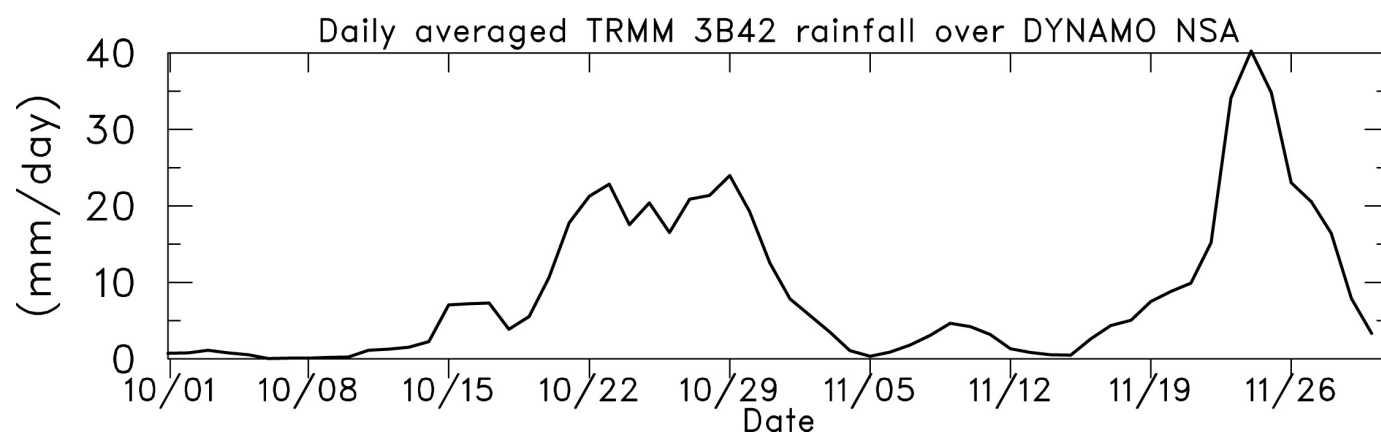
Latent and radiative heating anomalies



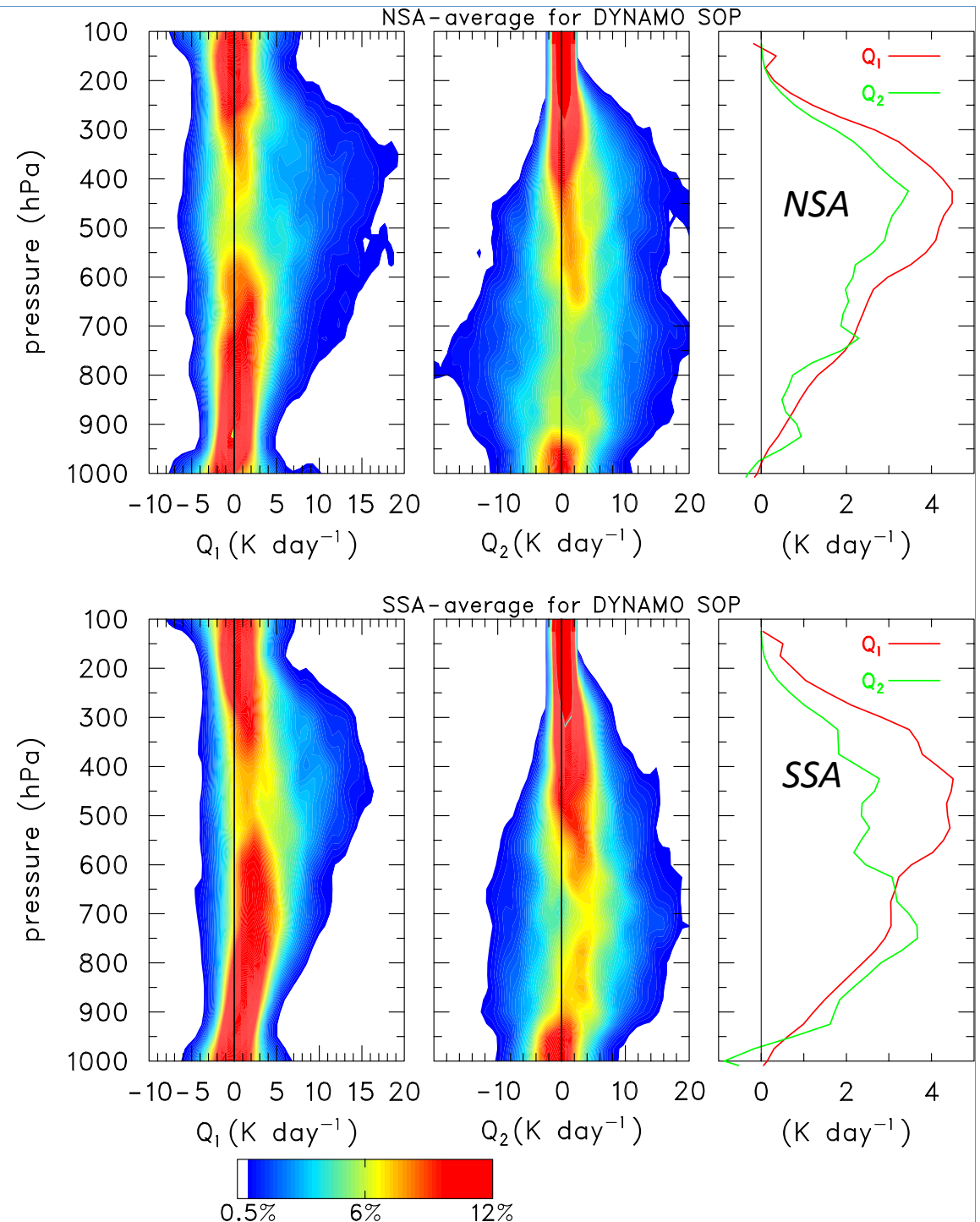
Ratio of cloud radiative forcing to convective heating

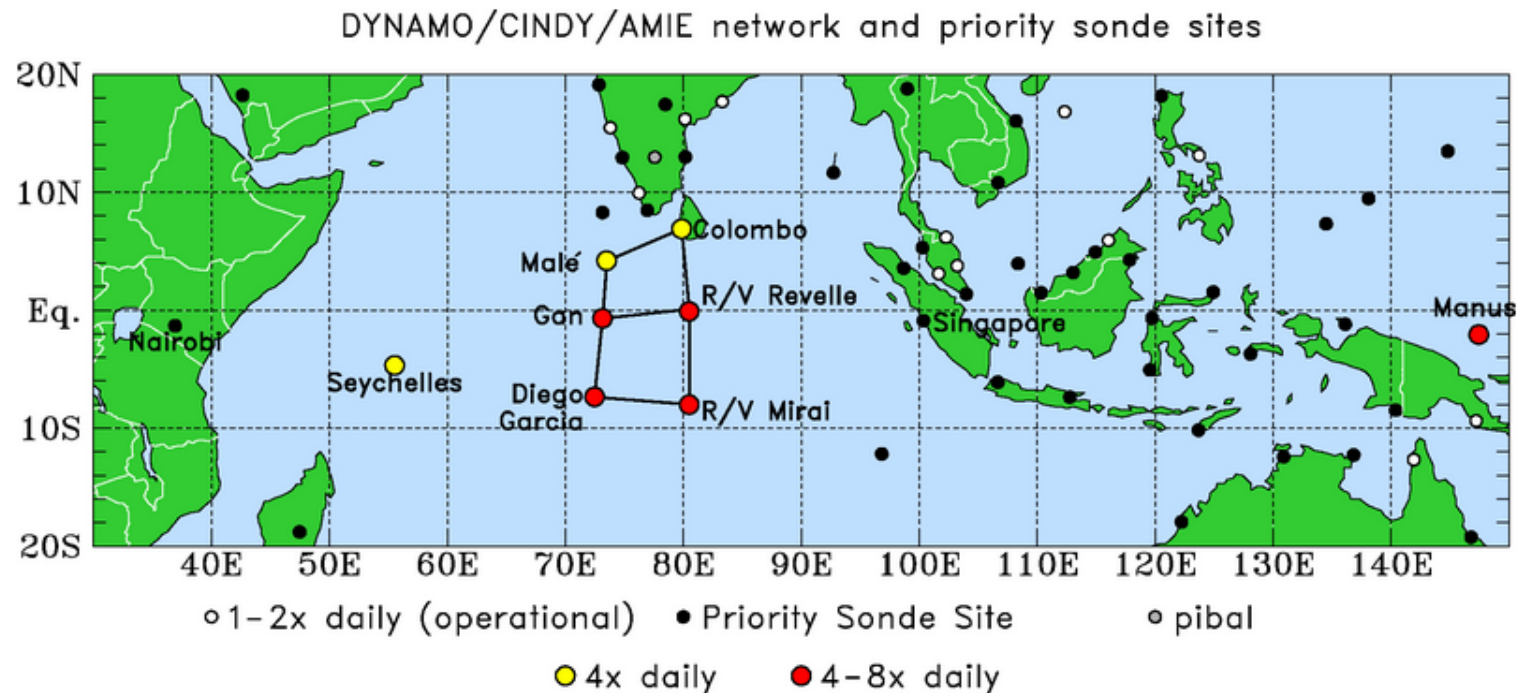


TRMM rainfall



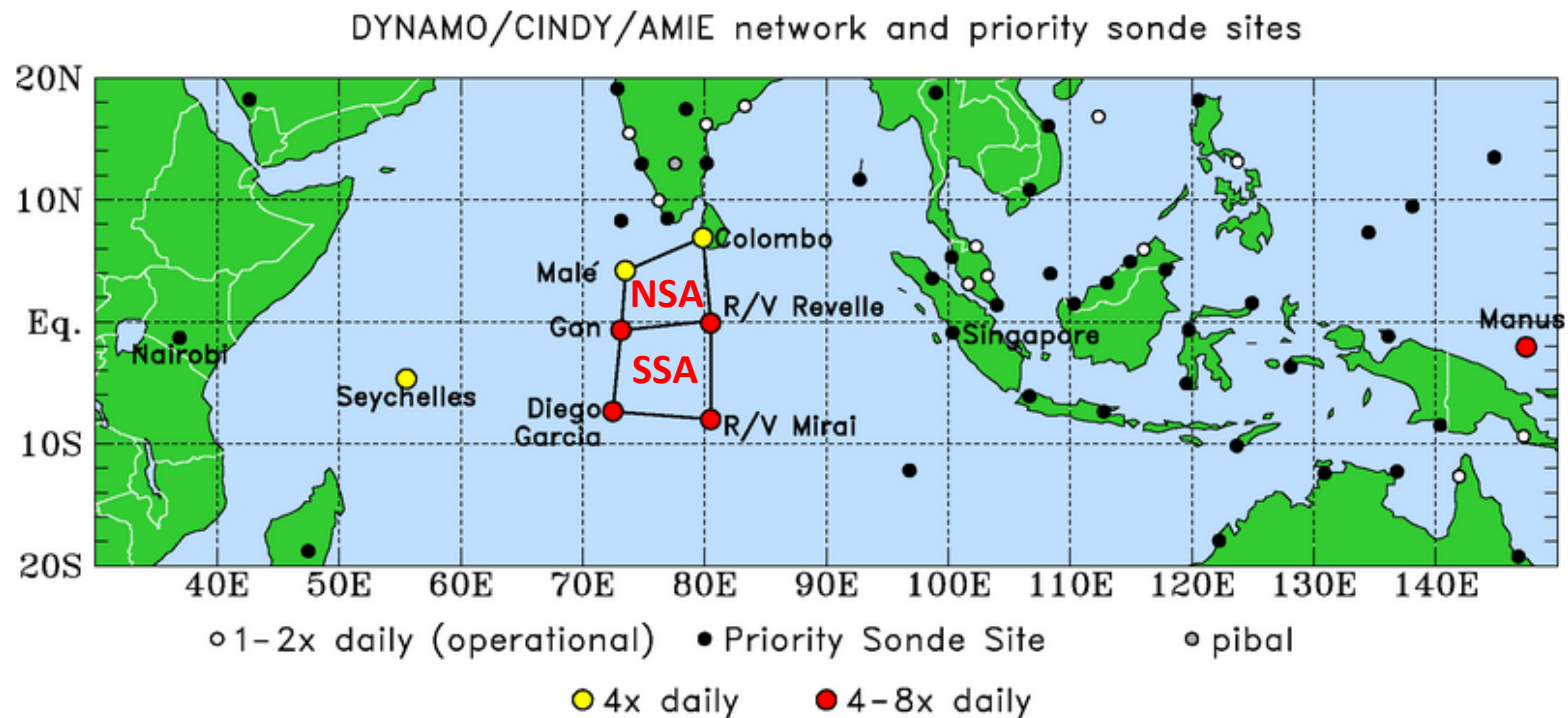
- ◆ Mean profiles of Q_1 and Q_2 for NSA and SSA differ
- ◆ NSA more stratiform; SSA more convective – supported by TRMM measurements (SF = 55% for NSA, 50% for SSA)
- ◆ Consistent with findings by Lin et al. (2004) that MJO has greater stratiform rain fraction than tropical mean





Input for Gridded Analyses

- Level 4 sounding data (uniform 5-hPa resolution) augmented with:
 - CIMSS cloud-drift satellite winds, ASCAT surface winds, and COSMIC thermodynamic profiles (T_d not used below 850 hPa)
- Enhancements to sounding data set:
 1. To better resolve the diurnal cycle, the 4/day soundings at Malé and Colombo were interpolated to a 3-h time resolution.
 2. An adjustment procedure was developed, making use of low-level ECMWF OA, to mitigate the Sri-Lanka island effects on Colombo soundings (Ciesielski et al. 2014)



Details for Objectively Analyzed Gridded Product

- Multiquadric interpolation to produce gridded analyses of basic fields
 - 3-hr, 1° horizontal res. (35-155°E, 20°S-20°N), 25-hPa vertical res. (sfc to 50 hPa)
 - Two versions of analyses were created: (V1) observations only, (V2) observations supplemented with ECMWF Operational Analyses (OA) in data sparse regions *outside* of core sonde network.
- Two versions produce similar results; except V2 provides a modest improvement in the budgets when ships are offsite
- Diagnostic fields computed with V2 analyses; results shown as array averages for the SOP (1 Oct. – 30 Nov. 2011) when sonde network was most complete.

Evaluating budgets through integral constraints

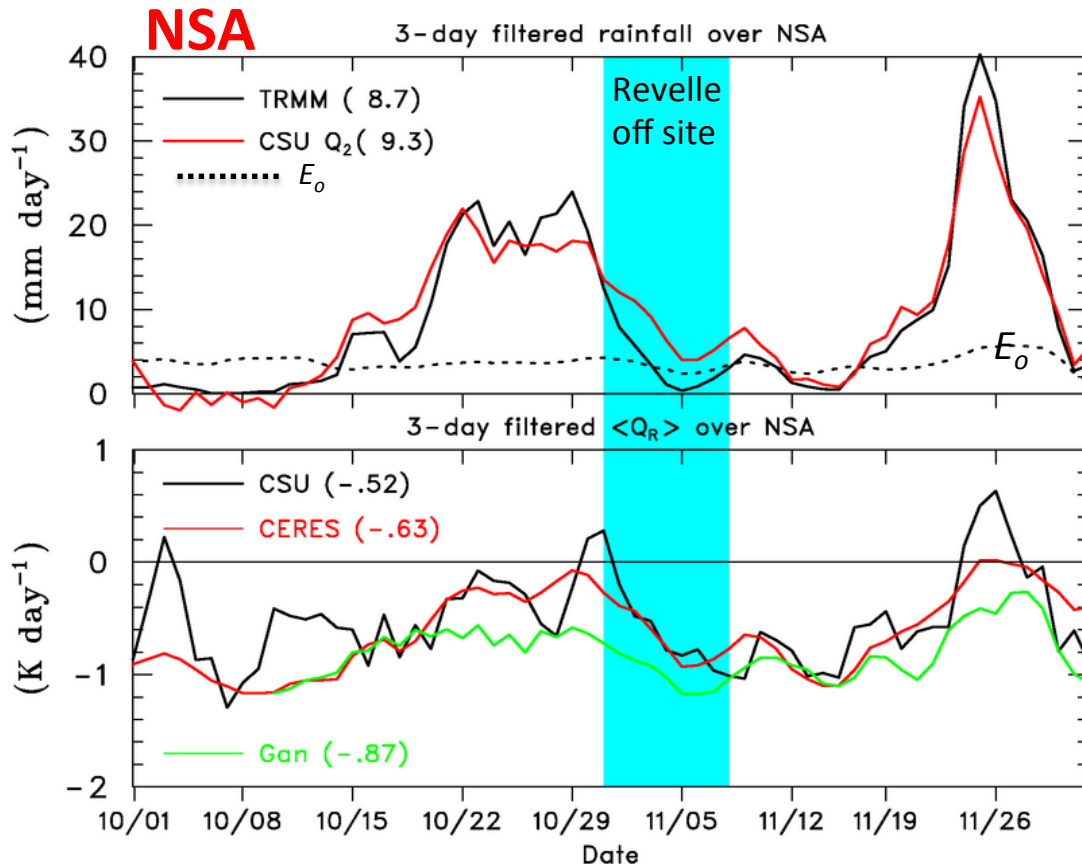
$$P_o = \frac{1}{L_v} \int_{P_T}^{P_S} Q_2 \partial p + E_o$$

E_o, S_o : WHOI OAFlux product (1° , daily)

$$\langle Q_R \rangle = \int_{P_T}^{P_S} (Q_1 - Q_2) \partial p - S_o - L_v E_o$$

Validated against:

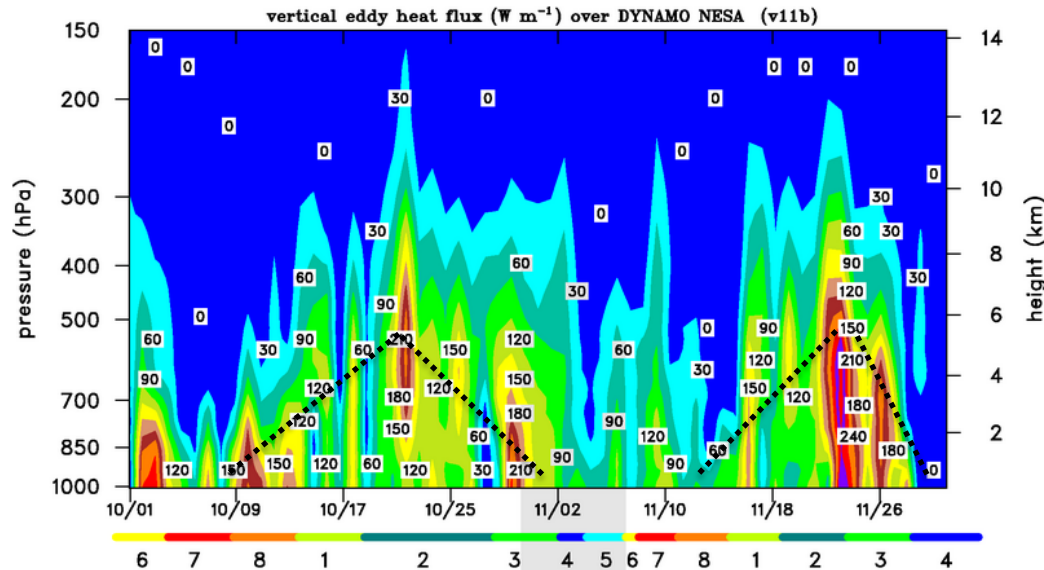
- TRMM 3B42V7 rainfall (3hr, 0.25°)
- CERES $\langle Q_R \rangle$ SYN1d product (3hr, 1°)
- Gan $Q_R(p)$, ARM CombRet (30s, 238 levels)



- Overall, excellent agreement between budget and independent estimates
- Largest P_o errors occur when Revelle is off site; budget uncertainty is large
- Other periods of P_o disagreement may be related to cloud-storage of water vapor.
- $\langle Q_R \rangle$ strongly modulated on MJO time scale, with much-reduced cooling during the active phases

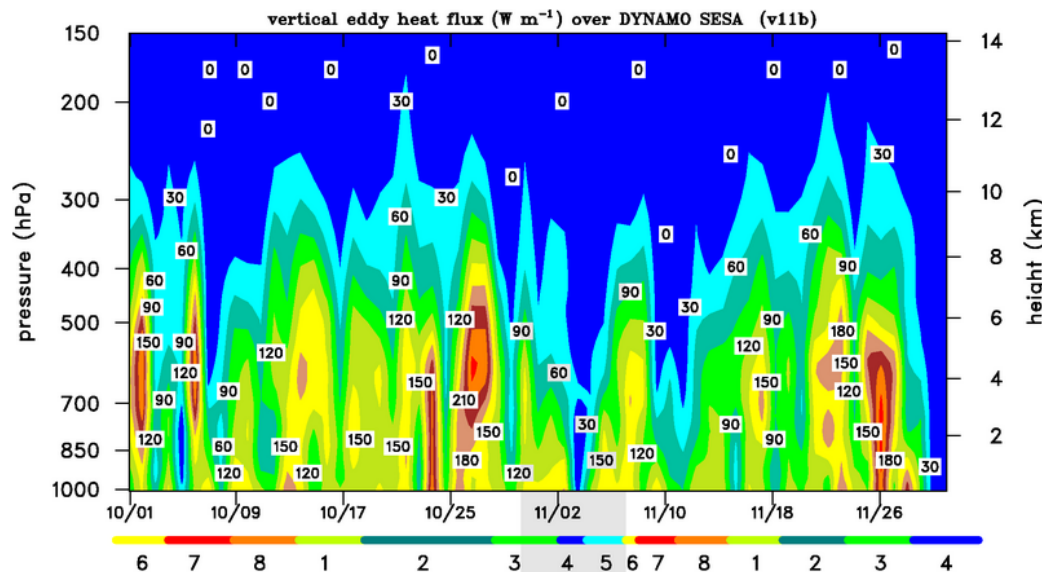
Time series of eddy flux of moist static energy $F(p)$

NSA



- Over NSA, eddy fluxes are strongly modulated on MJO time scale – being quite shallow during suppressed periods then deepening to peak at mid-level and gradually weakening near end of active phase.

SSA



- SSA is characterized with more persistent, deeper eddy fluxes.

Comparison of DYNAMO to TOGA COARE

| Array | Size | # of sonde sites | Duration | Sounding frequency |
|---------|-----------------------|------------------|----------|--------------------|
| DYN-NSA | (707 km) ² | 4 | 61 days | 4-8/day |
| DYN-SSA | (830 km) ² | 4 | 61 days | 8/day |
| TC IFA | (474 km) ² | 6 | 120 days | 4/day |

- NSA size is 50% > TC IFA
- SSA size is 75% > TC IFA

Experiment Mean Rainfall (mm day⁻¹)

| Array | Independent estimate | Q ₂ Budget | <i>r</i> |
|------------------|----------------------|-----------------------|----------|
| DYNAMO NSA | 8.7 | 9.3 | 0.96 |
| DYNAMO SSA | 9.0 | 10.4 | 0.91 |
| TOGA COARE (IFA) | 8.3 to 9.3 | 8.4 | 0.84 |

- High temporal correlations (*r*) indicate that budgets are accurately capturing the large-scale forcing signal.

Experiment mean $\langle Q_R \rangle$ (K day⁻¹)

| Array | Independent estimate | Budget | <i>r</i> |
|------------------|----------------------|--------|----------|
| DYNAMO NSA | -0.63 | -0.52 | 0.68 |
| DYNAMO SSA | -0.62 | -0.44 | 0.73 |
| TOGA COARE (IFA) | -0.38 to -0.84 | -0.55 | 0.43 |

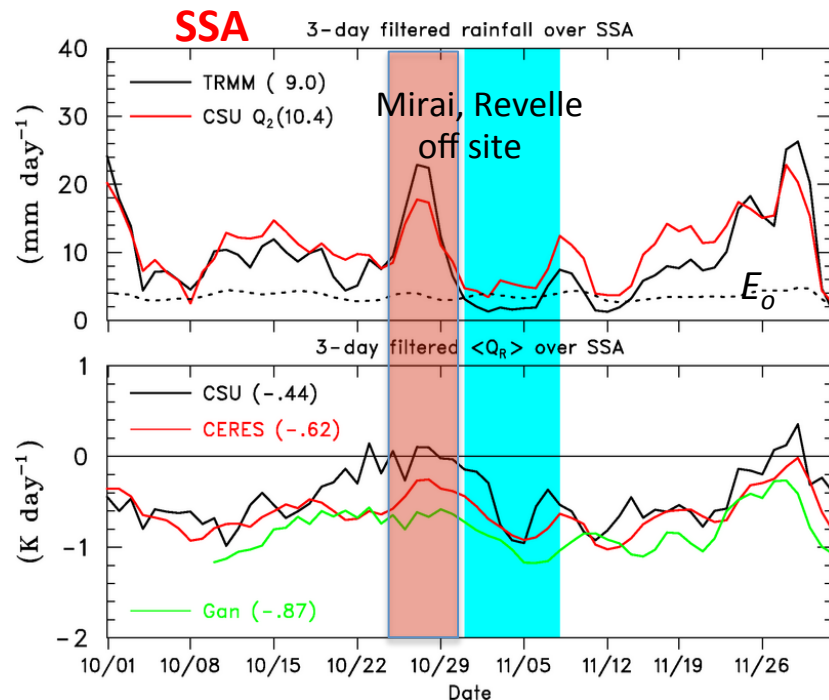
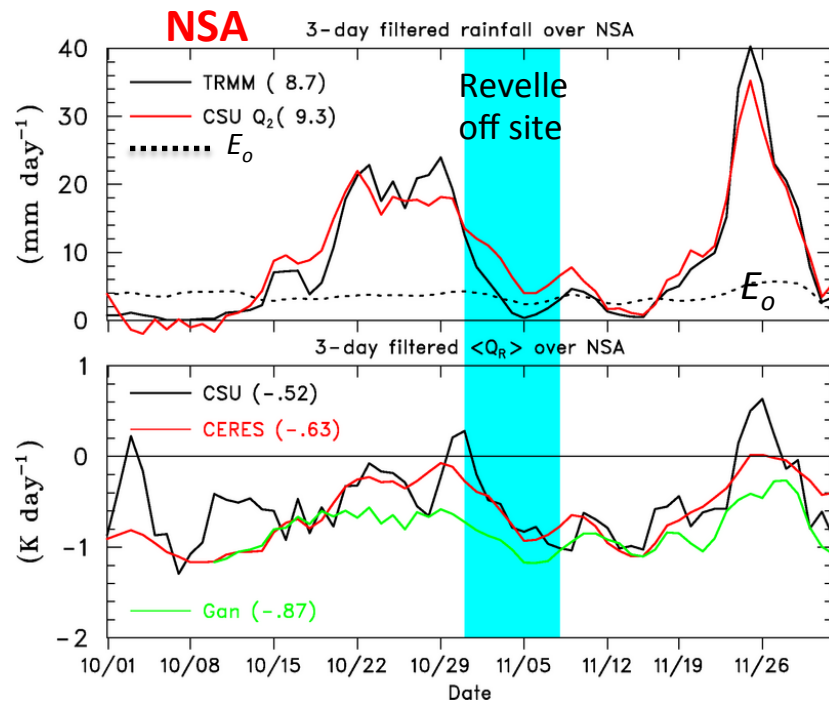
- Budget-estimated mean rainfall in DYNAMO is slightly higher than in TC and $\langle Q_R \rangle$ slightly less cooling

Supplemental material

Summary

- Budget analyses are based on a 3-h, 1°-gridded dataset which used high-quality upper-air soundings, satellite winds and COSMIC thermodynamic profiles as input.
- Supplementing these observation with model analyses results in a beneficial impact to the budgets when the ships are off site.
- Good agreement of budget-derived integral constraints (i.e. rainfall, $\langle Q_R \rangle$, surface fluxes) with independent estimates lends confidence that **the budgets are accurately capturing the large-scale forcing signal.**
- Convection over NSA is strongly modulated by MJO signal, while the SSA experienced more persistent, briefer episodes of rain related to ITCZ convection.
- Weaker vertical eddy fluxes over the NSA compared to SSA suggest a higher stratiform rainfall fraction over the NSA (Lin et al. 2004).
- Plans are underway to create an ensemble forcing dataset such that the effects of random sampling errors on budgets can be investigated.

Thank you



$$P_o = \frac{1}{L_v} \int_{P_T}^{P_S} Q_2 \partial p + E_o$$

$$\langle Q_R \rangle = \int_{P_T}^{P_S} (Q_1 - Q_2) \partial p - S_o - L_v E_o$$

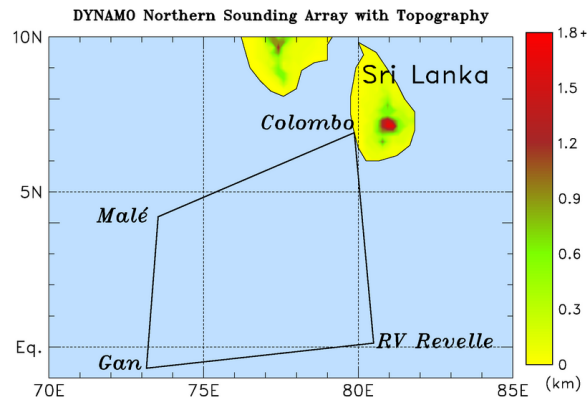
E_o, S_o : WHOI OAFlux product (1°, daily)

Validated against:

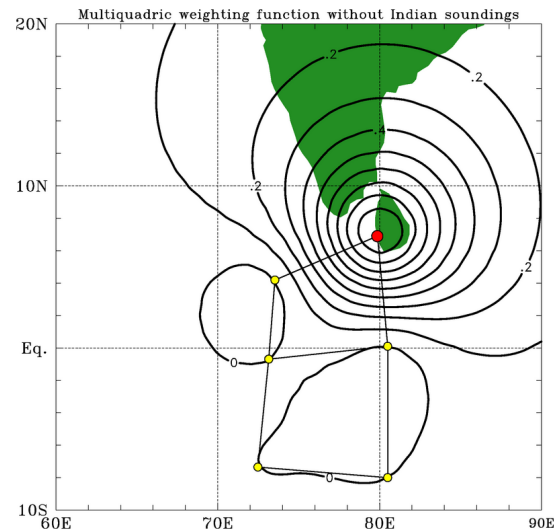
- TRMM 3B42V7 rainfall (3hr, 0.25°)
- CERES <Q_R> SYN1d product (3hr, 1°)
- Gan Q_R(p), ARM CombRet (30s, 238 levels)

- Rainfall over SSA is more persistent and episodic than over NSA
- Correlations lower than over NSA
- Revelle off site 29 Oct. to 09 Nov.
- Mirai off site 24 Oct. to 30 Oct.

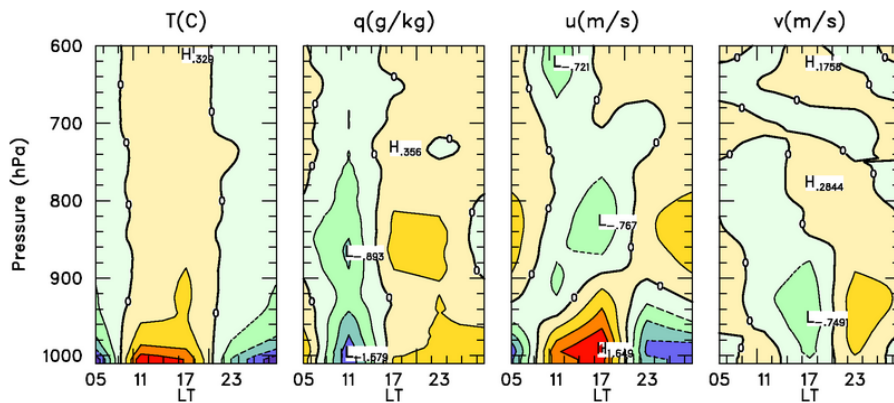
Mitigation of Sri Lanka Island effects on Colombo soundings (Ciesielski et al. 2014)



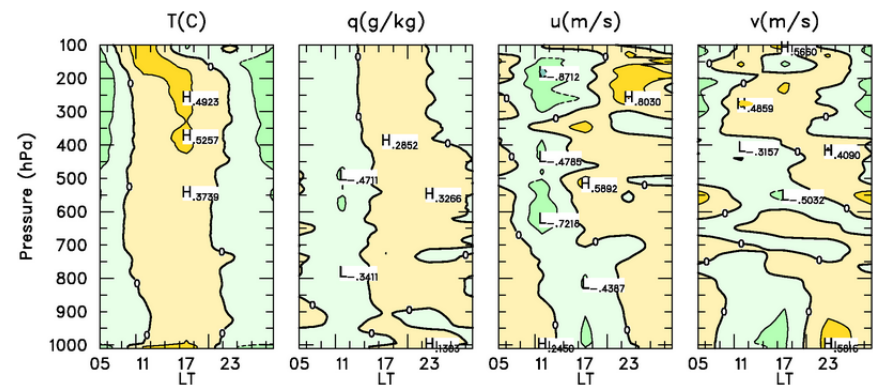
- Peak elevation east of Colombo was ~2500 m



- MQD weighting function showing how the influence of the Colombo data spreads horizontally

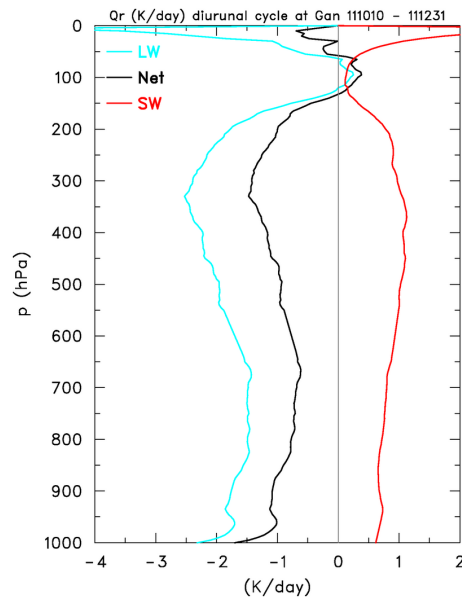


Diurnal cycle anomaly of original Colombo sounding fields.

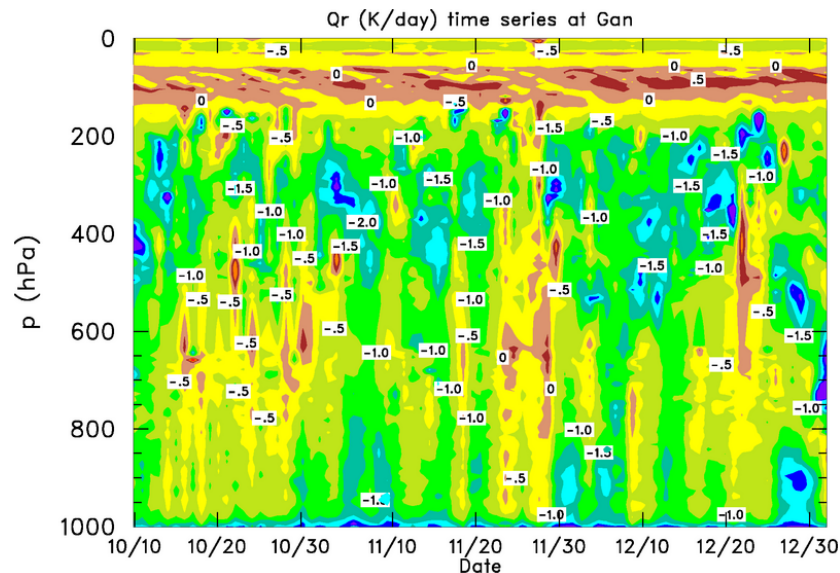
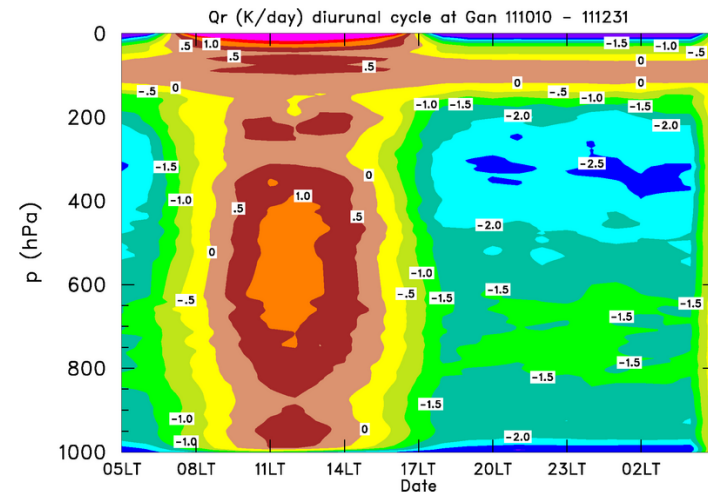


Diurnal cycle anomaly of adjusted Colombo sounding fields; DC at low levels has been effectively muted.

CombRet (Combined Retrieval) – radar-lidar cloud retrieval with radiative transfer model (Feng et al. 2013)



- Period-mean vertical profiles of LW, SW and Q_R (values below melting have been interpolated over)
- Period mean diurnal cycle of Q_R



- Height-time plot of Q_R at Gan
- To fill in 1 – 9 Oct. period, we used the mean diurnal cycle for the period 12-17 Nov. which had similar rainfall and budget profiles to early October period.

CERES (Clouds and the Earth's Radiant Energy System)

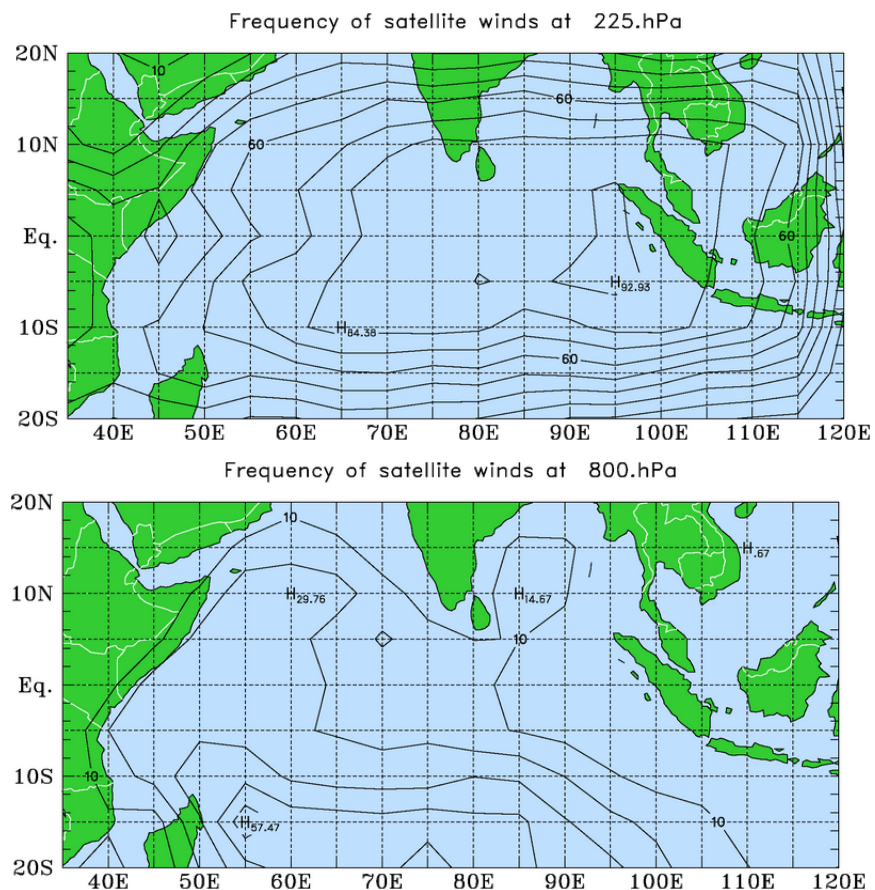
- SYN1deg 3-hr Ed3A Product (Wielicki et al. 1996)

Input: Terra, Aqua, MODIS and GEO (MET-7 at 63E)

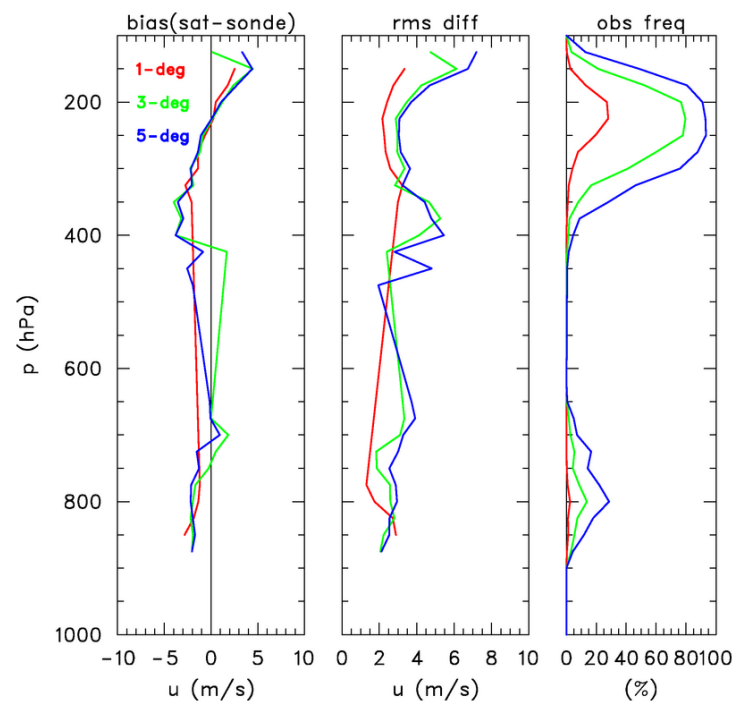
Products: TOA and surface radiative fluxes at 1°, 3-hr resolution

Fluxes are produced using Langley Fu-Liou radiative transfer model

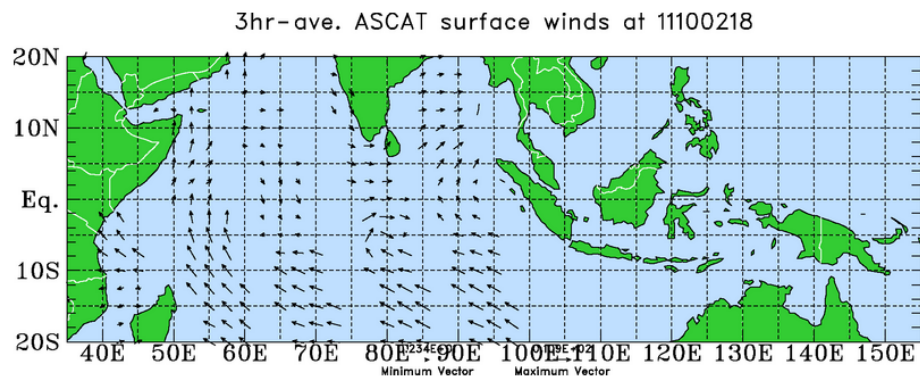
Coverage of CIMSS satellite winds



Over central IO CIMSS winds are available about ~80% of time at upper level and 20% at 800 hPa (assuming 3-h, 5° resolution and ≥ 3 obs in average)

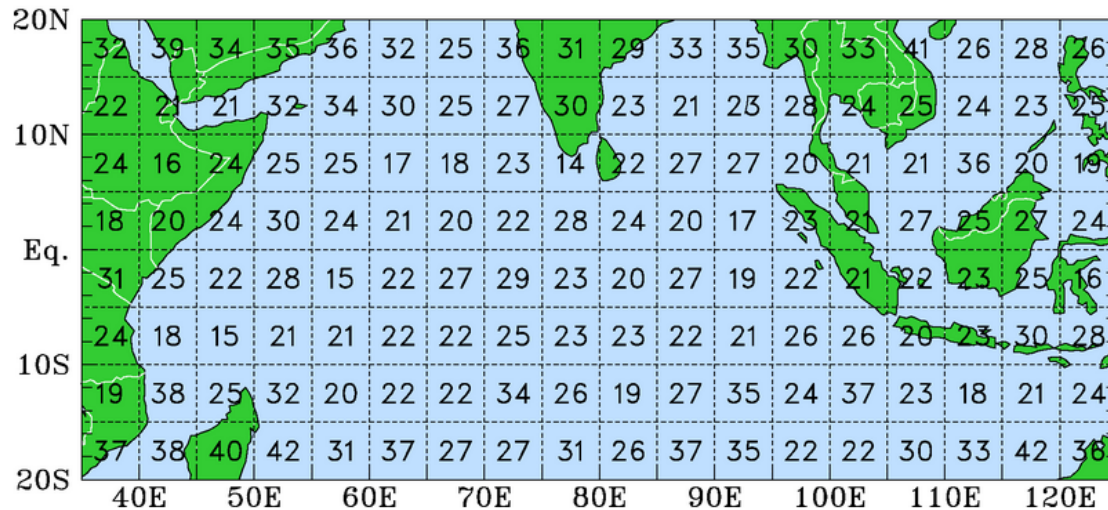


Example of data coverage of ASCAT surface winds for 3 hour window. At a given point ASCAT winds are available ~15% of the time assuming at 3-h resolution.

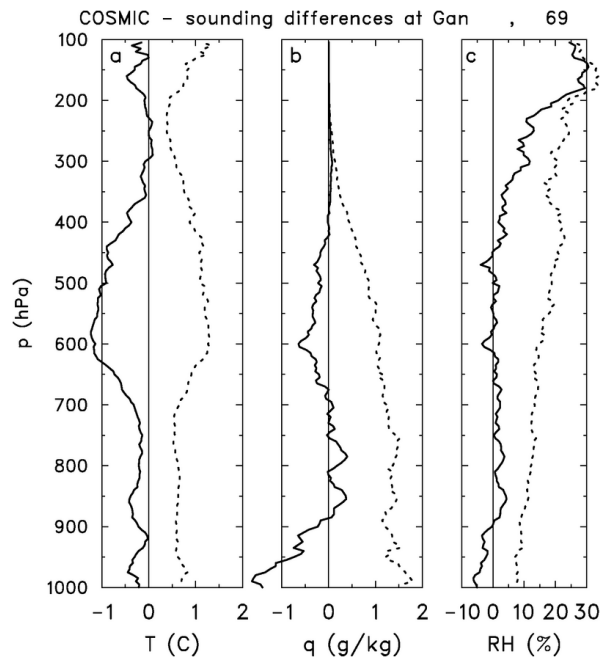


COSMIC data

COSMIC inventory for 01 Oct. – 31 Dec. 2011



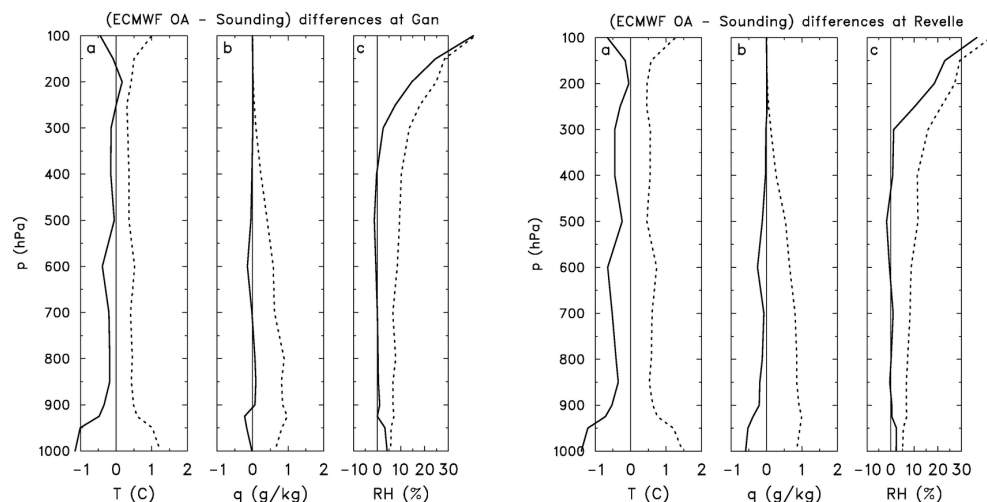
- Number of COSMIC soundings in a 5°x5° box for 3 month period
- Averages ~ 1 sounding every 3-4 days



- Biases and RMS differences at Gan between COSMIC data and Gan soundings using COSMIC profiles within 5° radius of Gan
- COSMIC moisture retrievals not used below 850 hPa (issues with penetration of signal under moist conditions.)
- COSMIC shows 1°C cool bias at near 0°C level
- COSMIC upper-level RH bias due to ECMWF moisture bias at these levels.

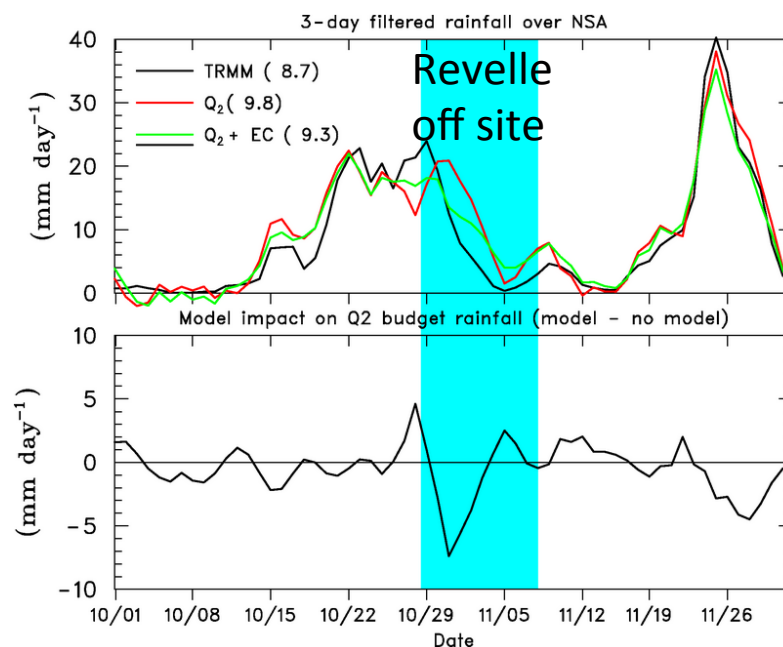
ECMWF Operational Analyses

- 6h, 0.25° resolution, 20 pressure levels (sfc to 20 hPa)
- 95% of soundings from core array reached operational centers



- ECMWF model biases and RMS differences from collated soundings
- Model has cool, dry bias at low-levels and moist bias above 250 hPa.

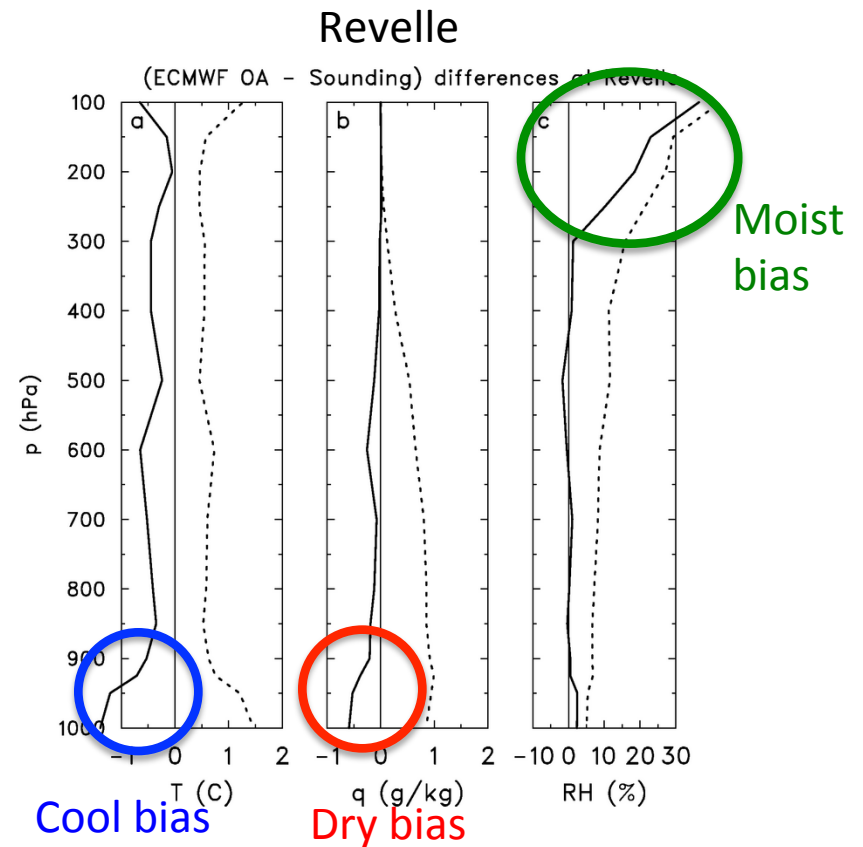
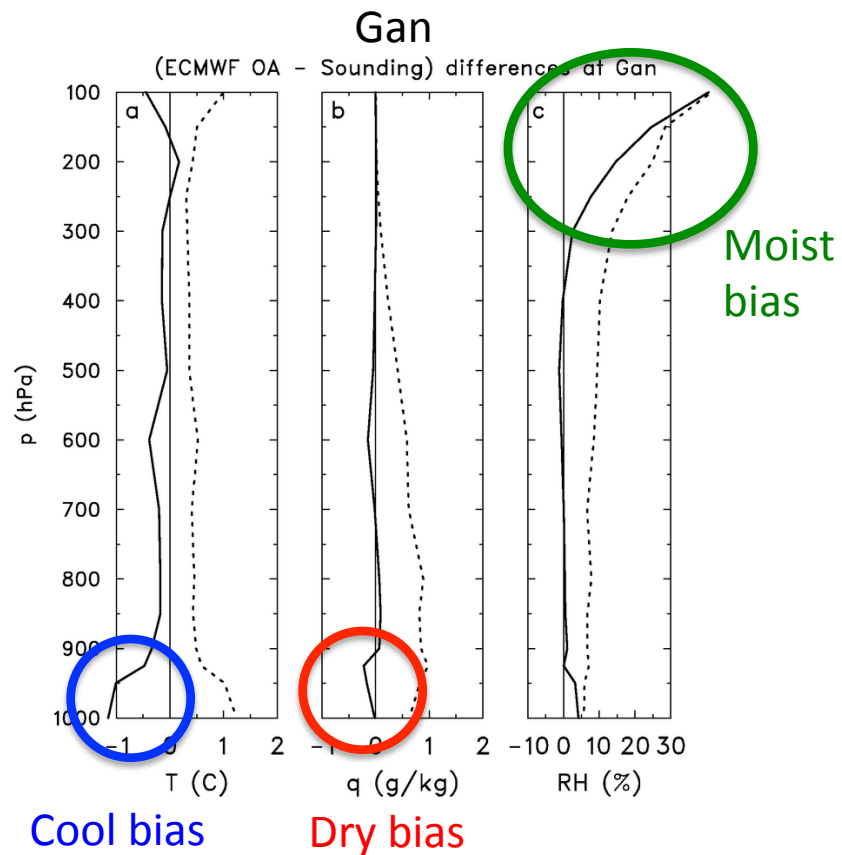
- Using model data to supplement observations has largest impact when ships were off site.



ECMWF Operational Analyses

- 6h, 0.25° resolution, 20 pressure levels (sfc to 20 hPa)
- 95% of soundings from core array reached operational centers
- Comparison of model and sonde data show the model analyses have a slight low-level cool and dry bias and an upper-level moist bias (too much cirrus)
- The main impact of using ECMWF analyses is seen during periods when ships were offsite

Bias (solid curve), RMS difference (dashed)



Heat and Moisture Budgets

Apparent heat source/moisture sink:

(Yanai et al. 1973)

$$Q_1 \equiv c_p [(\partial \bar{T} / \partial t + \bar{\mathbf{v}} \cdot \nabla \bar{T} + (p/p_0)^\kappa \bar{\omega} \partial \bar{\theta} / \partial p)]$$

$$Q_2 \equiv -L(\partial \bar{q} / \partial t + \bar{\mathbf{v}} \cdot \nabla \bar{q} + \bar{\omega} \partial \bar{q} / \partial p)$$

Integrated budgets:

[Definition: $\langle \quad \rangle \equiv 1/g \int_{p_T}^{p_a} (\quad) dp$]

$$\text{Heat : } \langle Q_1 \rangle = \langle Q_R \rangle + LP + S \quad (1)$$

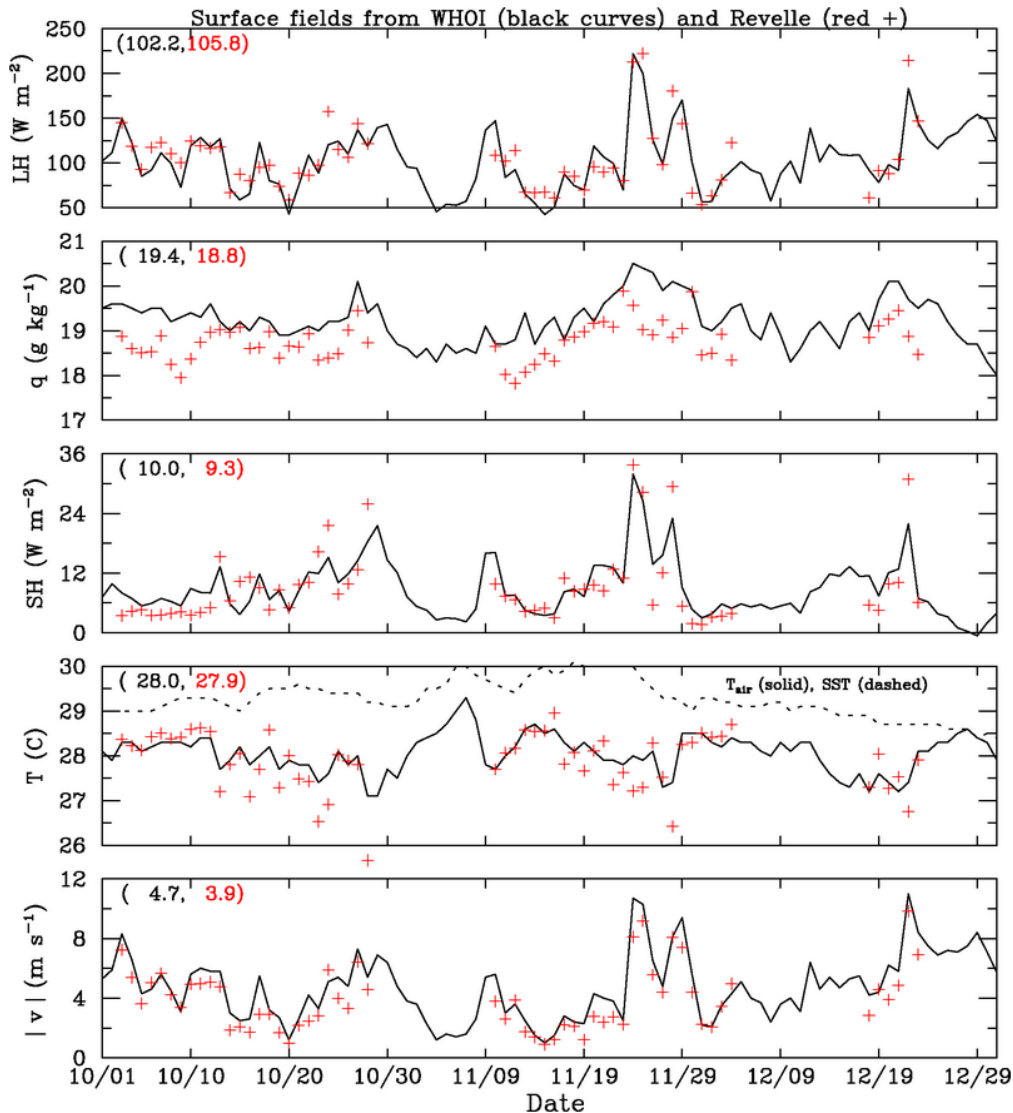
$$\text{Moisture : } \langle Q_2 \rangle = L(P - E) \quad (2)$$

(1) – (2) yields

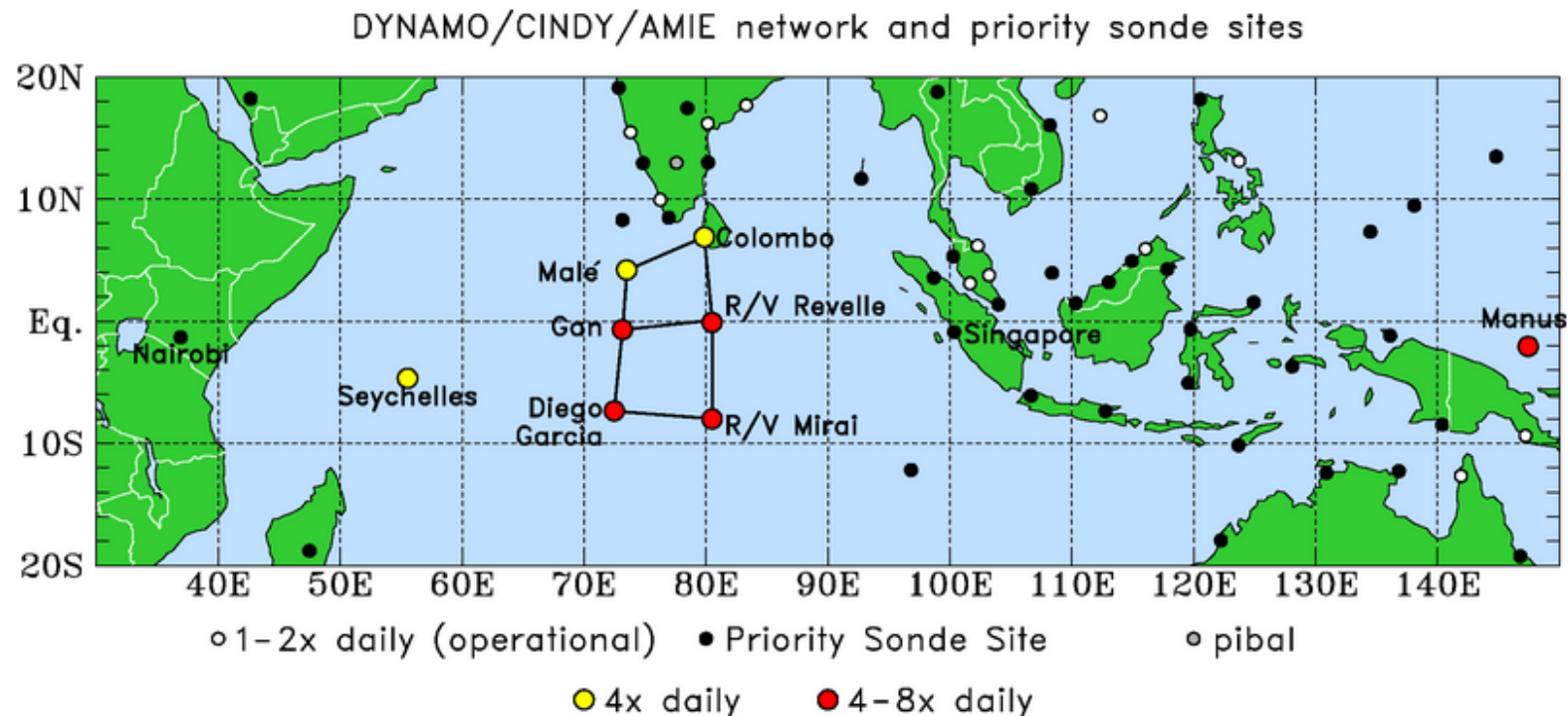
$$\langle Q_1 \rangle - \langle Q_2 \rangle = \langle Q_R \rangle + S + LE, \quad (3)$$

where P = Precipitation rate, E = Evaporation rate, S = Sensible heat flux

Comparison of fluxes at Revelle (in situ vs WHOI estimated)



- Period mean LH flux at Revelle LH fluxes is a few percent higher than WHOI
- Surface q from Revelle slightly lower than WHOI values
- Period mean SH flux from Revelle $\sim 5\%$ less than WHOI flux.

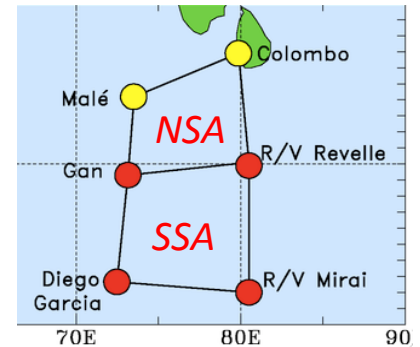


DYNAMO/AMIE/CINDY sounding dataset

- Consist of ~26,000 upper-air soundings; ~ half are high-vertical resolution (hi-res)
- All hi-res soundings were corrected, if necessary (Ciesielski et al. 2014)
- Comparisons with independent datasets (such as ground based GPS and MWR retrievals of PW) suggest that the sounding-based humidity observations are of very high quality.
- This study uses the Level 4 (uniform 5-hPa resolution with QC flags) soundings and will focus on the Special Observing Period (**SOP**) 1 Oct. – 30 Nov. 2011.

Enhancements to sounding dataset

1. To better resolve the diurnal cycle and take advantage of the 8/day sampling at sites in the southern array, the 4/day soundings at Malé and Colombo were interpolated to a 3-h time resolution.
 2. An adjustment procedure was developed, making use of low-level ECMWF OA, to mitigate the Sri-Lanka island effects (i.e., both flow blocking and the island-induced diurnal cycle) on Colombo soundings (Ciesielski et al. 2014).
- Adjusted Colombo soundings are more representative of open-ocean conditions.



Examples of flow blocking at Colombo (red dot) due to topography to the east of site.

